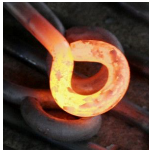


Generation of light

Thermal radiation and Luminescence

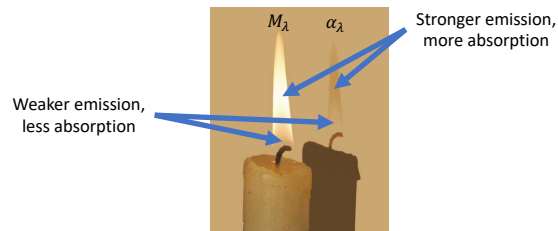


Zsolt Mártonfalvi

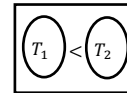
Kirchoff's law

Bodies that emit more also absorb more. The ratio between radiant emittance and absorption coefficient is constant within a narrow range of wavelength (λ):

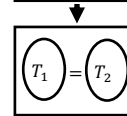
$$\frac{M_{\lambda \text{ body1}}}{\alpha_{\lambda \text{ body1}}} = \frac{M_{\lambda \text{ body2}}}{\alpha_{\lambda \text{ body2}}} = \text{constant}$$



Thermal radiation



The temperature of the two bodies will become equal with time even in vacuum!
(no convection or conduction)



Consequently:
All bodies emit radiation independently of the temperature of their surroundings. The emitted radiation is always electromagnetic radiation.

Quantitative description of thermal radiation:

- Radiant emittance (M)

$$M = \frac{\Delta P}{\Delta A} = \left[\frac{W}{m^2} \right]$$

- Absorption coefficient (α)

$$\alpha = \frac{I_{\text{absorbed by the surface}}}{I_{\text{received by the surface}}} \quad (0 \leq \alpha \leq 1)$$

M and α strongly depend on the absolute temperature of the body!

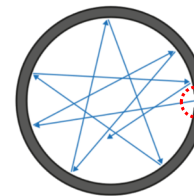
The ideal black body

A theoretical body that is used as a model in the description of thermal radiation. It absorbs all radiation that falls on it:

$$\alpha_{\text{black body}} = 1$$

thus we can calculate any real body's radiant emittance if we know its absorption coefficient:

$$M_{\lambda i} = \alpha_{\lambda i} M_{\text{black body}}$$



The hole on a dark cavity approximates an ideal black body

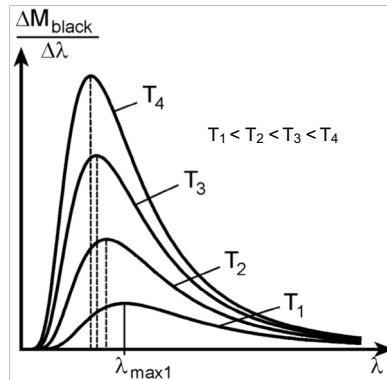
The radiant emittance of a black body depends strongly on the absolute temperature!

Stefan's law:

$$M = \sigma T^4 \quad \sigma = 5.67 \cdot 10^{-8} \frac{W}{m^2 K^4}$$

The ideal black body

The emission spectrum of the black body is **continuous**.



Wien's displacement law:

$$\lambda_{max} T = \text{constant}$$

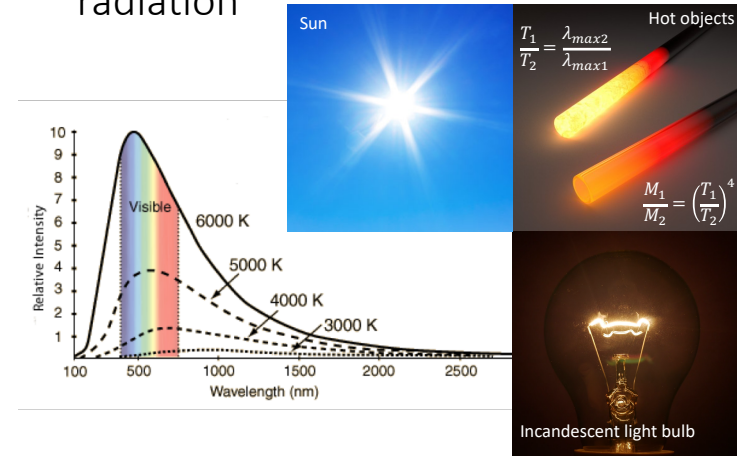
Max Planck's radiation law:

The radiation energy of a black body is emitted in discrete integer multiples of a fundamental "package", the quantum.

$$E = hf$$

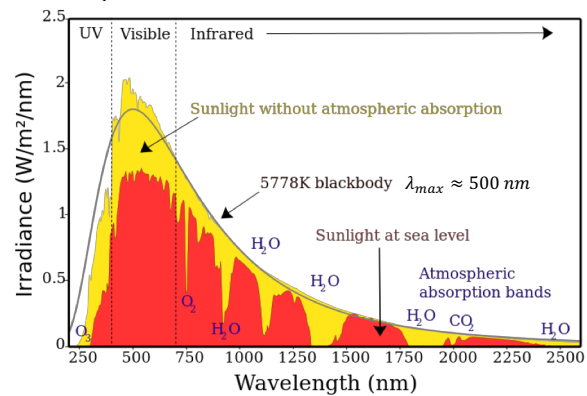
$$h = 6.62 \cdot 10^{-34} \text{ Js}$$

Light sources based on thermal radiation



Sun light

Spectrum of Solar Radiation (Earth)



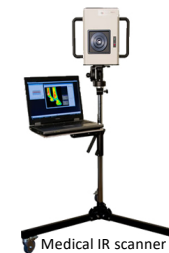
Medical applications

Telethermography

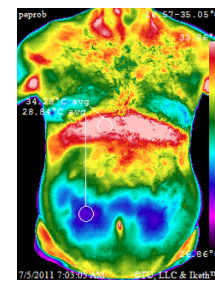
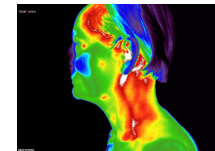
Measures the radiant emittance of thermal radiation of the human body.

Human body: $\lambda_{max} \approx 10 \mu\text{m}$ (infrared)
 $\alpha \approx 0.95$ (95% of ideal black body)

Radiant emittance strongly depends on temperature!



Diagnosis of inflammations, cancer, circulatory defects that cause the change in local temperature.



Luminescence

Spontaneous emission of a photon due to the relaxation of an excited electron.

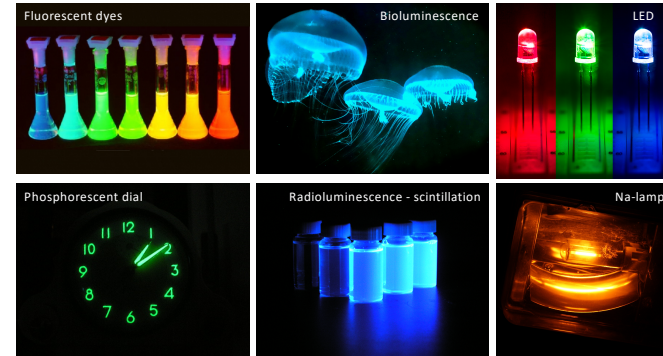
Phases of luminescence:

- Absorption of external energy
- Excitation
- Emission of energy in the form of electromagnetic radiation

Types of luminescence

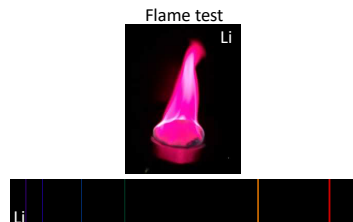
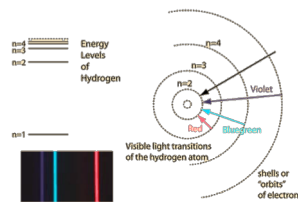
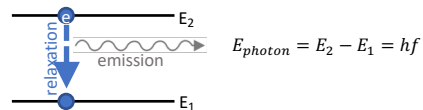
Type of excitation	Name	Example
Light	Photoluminescence	Fluorescent lamp
Electric	Electroluminescence	Hg vapor lamp
Radioactive	Radioluminescence	Nal (Tl)
Mechanical	Triboluminescence	Sugar cube
Biochemical	Bioluminescence	firefly
Thermal	Thermoluminescence	CaSO ₄ (Dy) (dosimeter)

Type of relaxation	Name	Example
S ₁ – S ₀ (fast)	Fluorescence	Fluorescein
T ₁ – S ₀ (slow)	Phosphorescence	Phosphorous



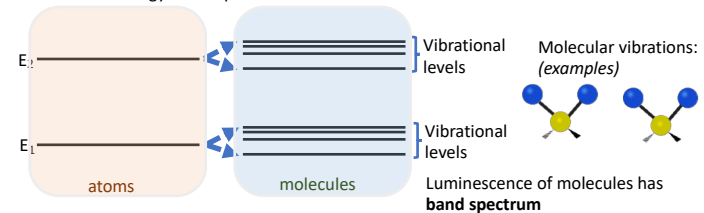
Luminescence of atoms

Luminescence emission of atoms has a **line spectrum**.





Luminescence of molecules

Discrete energy levels split into **vibrational levels**.



Spin states of excited electrons

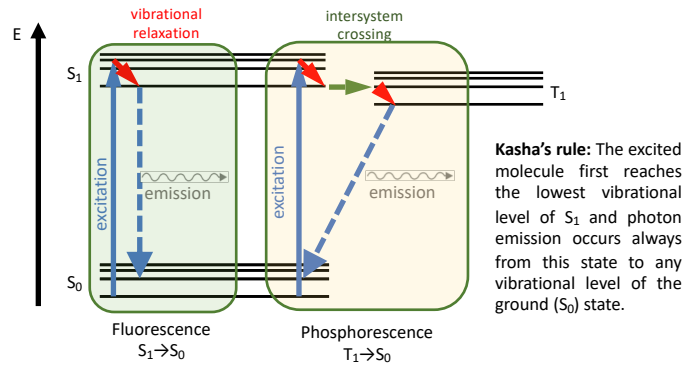
Singlet state (S) 
Sum of spin quantum numbers (S) is zero (+1/2, -1/2)

Triplet state (T) 
Sum of spin quantum numbers (S) is one (parallel spins)

Magnetic moment of spin state: $2S + 1$

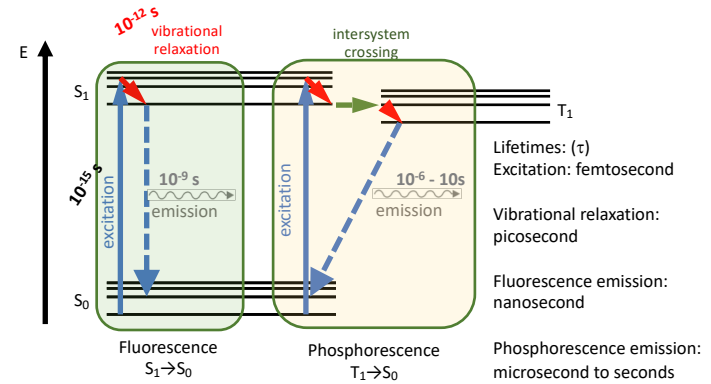
Luminescence emission of molecules

Jablonsky diagram



Luminescence emission of molecules

Jablonsky diagram



Luminescence emission of molecules

Quantum yield (Q_F):

$$Q_F = \frac{k_f}{k_f + k_{nr}} = \frac{\text{number of photons emitted}}{\text{number of photons absorbed}}$$

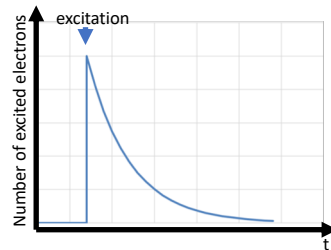
Lifetime (τ):

$$N = N_0 e^{-\frac{t}{\tau}}$$

$$\tau = \frac{1}{k_f + k_{nr}}$$

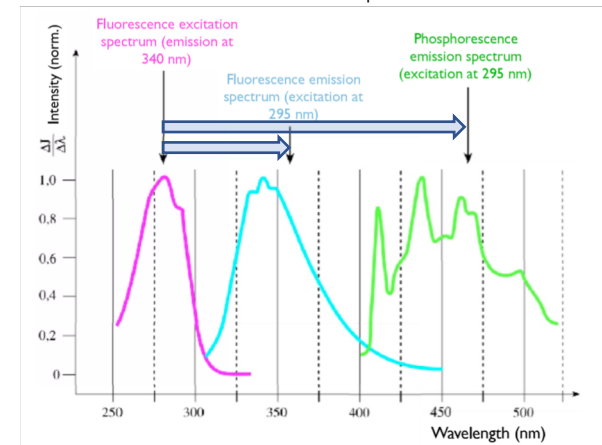
k_f : Rate of photon producing transitions

k_{nr} : Rate of non-radiative transitions



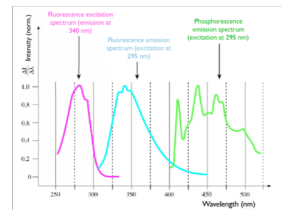
Stokes shift

The shift between the emission and excitation spectra



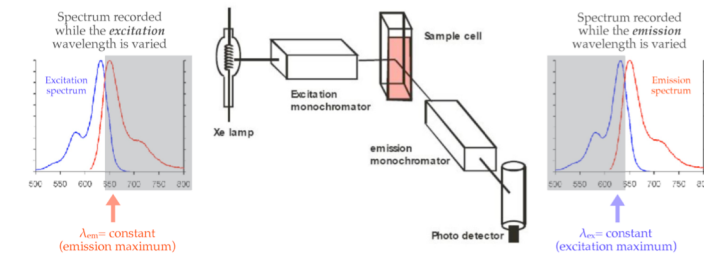
Comparing emission processes

	Fluorescence	Phosphorescence
Relaxation	From singlet state $S_1 \rightarrow S_0$	From triplet state $T_1 \rightarrow S_0$
Lifetime	Nanoseconds	From microseconds up to seconds
Stokes shift	Smaller	Larger (because T_1 is at lower energy than S_1)



Measurement of luminescence

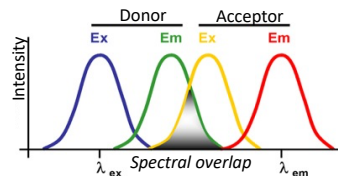
Luminescence spectrofluorimeter



Applications - FRET

Förster Resonance Energy Transfer

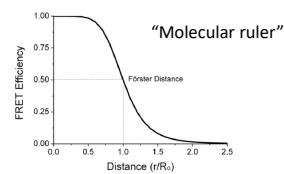
Energy transfers from donor without emission to acceptor in dipole-dipole interactions. Requires spectral overlap between donor emission and acceptor absorption.



FRET efficiency or quantum yield (E):

$$E = \frac{1}{1 + (r/r_0)^6}$$

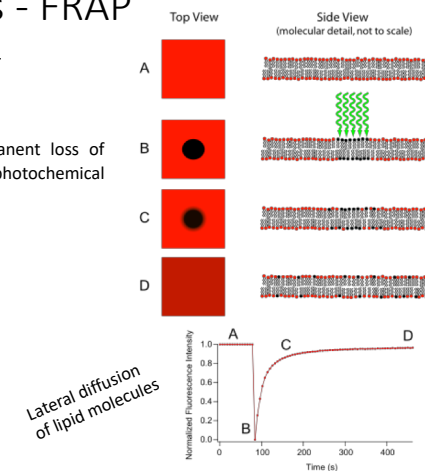
r_0 : Förster distance



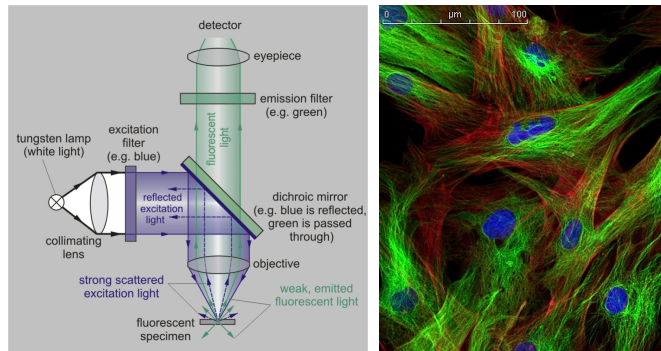
Applications - FRAP

Fluorescence Recovery After Photobleaching

Photobleaching: The permanent loss of fluorescence due to photochemical reactions.

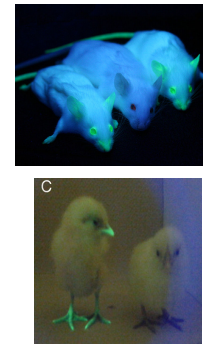


Fluorescence microscope



Biomedical applications

Fluorescent transgenic animals



Fluorescence guided surgery

