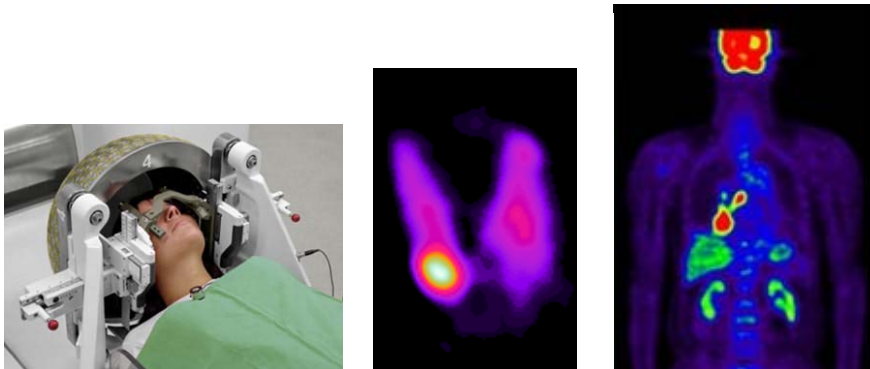


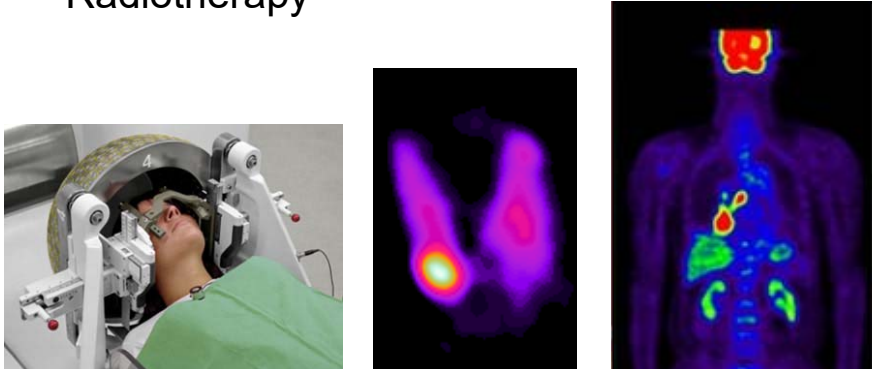
# The atomic nucleus. Radioactivity. Nuclear radiations

László Smeller









## Why?

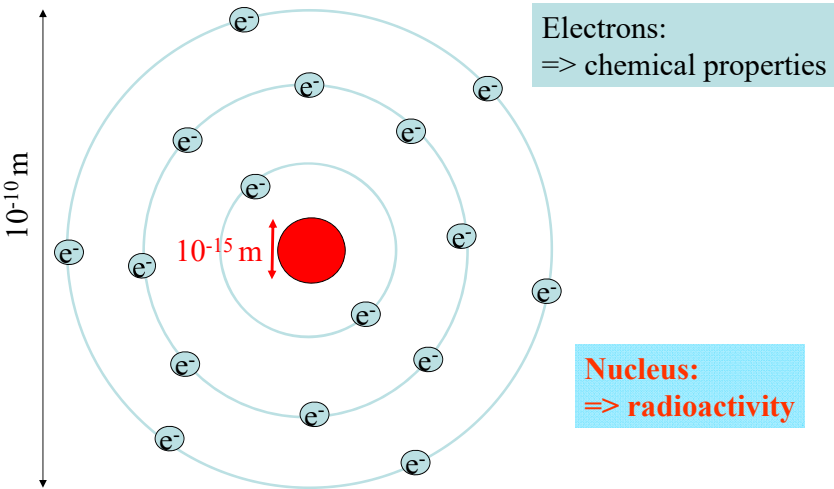
- Medical applications of the nuclear radiation:
- Nuclear imaging
  - Radiotherapy



## Length scale of the nature

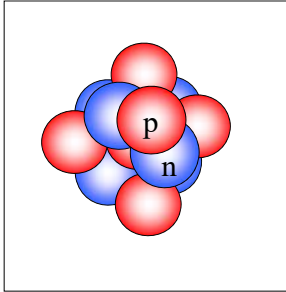
m		
10 <sup>0</sup>	meter	men 
10 <sup>-3</sup>	millimeter	size of letters you can read
10 <sup>-6</sup>	micrometer	size of a cell (e.g. erythrocyte)  Ø 7µm
10 <sup>-9</sup>	nanometer	protein 
10 <sup>-10</sup>	– angstrom	diameter of an atom, bond length H atom Ø ≈ 1 angstrom (Å) 
10 <sup>-12</sup>	picometer	wavelength of the X-ray 
10 <sup>-15</sup>	femtometer	size of the nucleus 

## The electrons and the nucleus



# Structure of the nucleus

Elementary charge =  $1.6 \cdot 10^{-19} \text{ C}$



	charge	mass
proton	+1 e	1 atomic mass unit
neutron	0	1 atomic mass unit

A (mass number) = number of protons + number of neutrons

Z (atomic number) = number of protons

99 nucleon, 43 proton and 56 neutron

99 Tc  
43

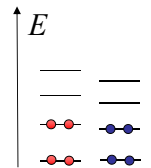
# Stability of the nucleus

- Coulomb force: destabilization  
(electrostatic repulsion between the protons)

- Nuclear force: very strong attractive force  
acts only on short range (~fm)  
independent on the charge



- Quantized energy levels for the nucleus.
- Typical binding energy is in the MeV range  
 $eV = 1.6 \cdot 10^{-19} \text{ J}$



# Isotopes

Variants of an element, where the

- number of protons are the same

=> the chemical properties are identical.

- number of neutrons are different

=> the atomic numbers are different



unstable  
(radioactive)

stable

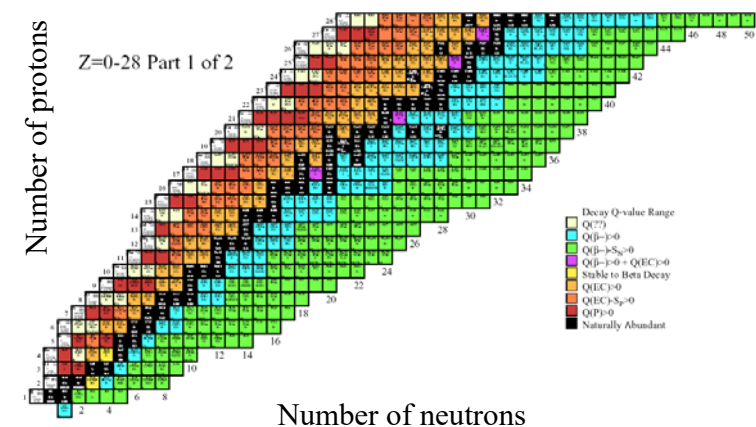
unstable  
(radioactive)

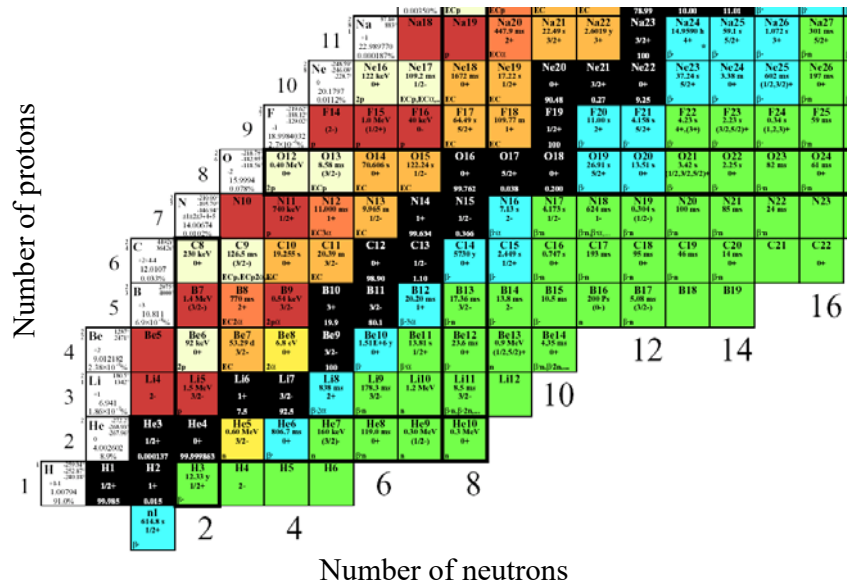
isotope <-> radioactive isotope



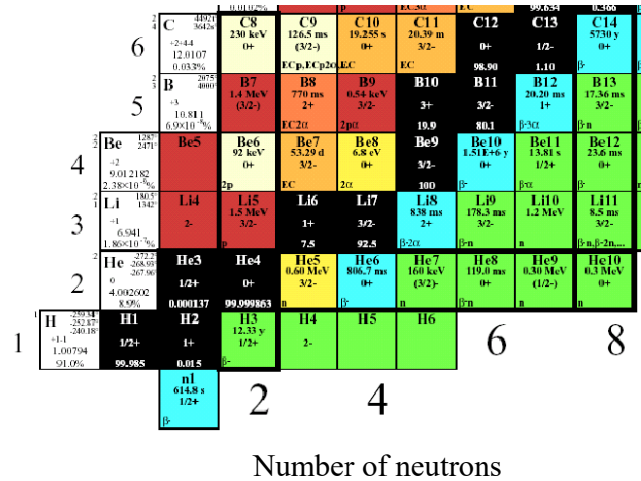
György Hevesi

# Table of isotopes





Number of protons



Number of neutrons

## Radioactive decays and particles

$\alpha$  - decay

$\alpha$  - particle =  ${}^4_2\text{He}$  nucleus

$\beta^-$  - decay :  $\beta^-$   
 $\beta^+$

$\beta^-$  particle = electron

$\beta^+$  particle = positron

Isomeric transition

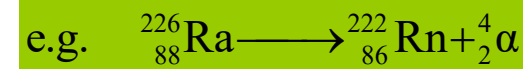
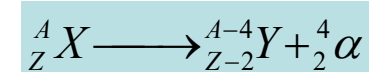
$\gamma$ -ray

K-electron capture

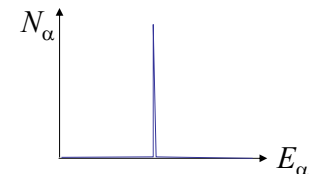
characteristic x-ray photon

## $\alpha$ - decay

$\alpha$  - decay: an  $\alpha$  particle ( ${}^4\text{He}$  nucleus) will be emitted  
typical for the heavy atoms



Line spectrum  
 $E_\alpha \sim \text{MeV}$



neutron surplus

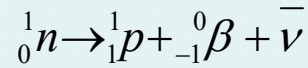
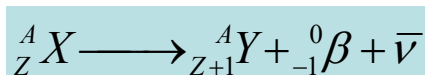
## $\beta^-$ - decay

e.g.:  $^{20}_{9}\text{F}$

$^{32}_{15}\text{P}$

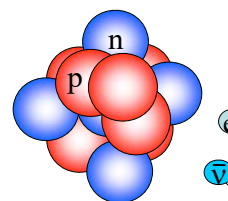
$^{59}_{26}\text{Fe}$

$^{131}_{53}\text{I}$



remains in  
the nucleus

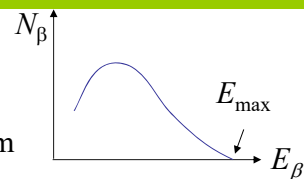
leave the  
nucleus



$\beta^-$ -ray

$$\beta^- = ^0_{-1}\beta = e^-$$

continuous  
energy spectrum



proton surplus

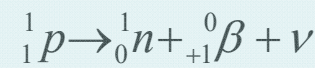
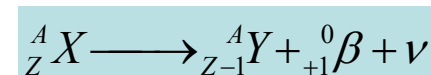
## $\beta^+$ - decay

e.g.:  $^{11}_6\text{C}$

$^{15}_8\text{O}$

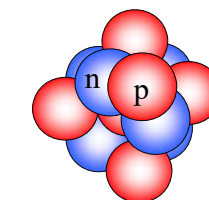
$^{18}_9\text{F}$

$^{52}_{26}\text{Fe}$



remains in  
the nucleus

leave the  
nucleus



$\beta^+$ -ray

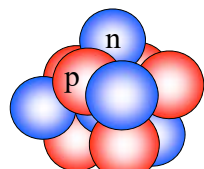
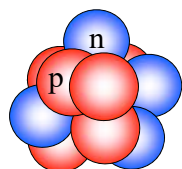
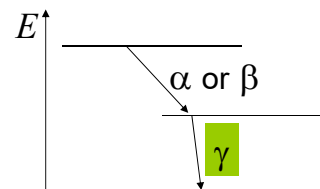
continuous energy spectrum

These isotopes must be produced  
artificially (e.g. in cyclotron)

## Prompt $\gamma$ -radiation

The daughter nucleus has an **energetically unfavoured** arrangement of nucleons.

The surplus energy will normally be emitted immediately (<ps) in form of the  $\gamma$  radiation



Atomic number, mass number are unchanged.

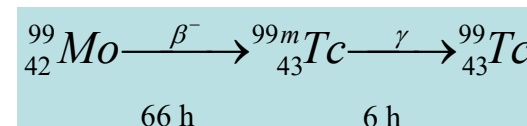
## Isomeric transition

If the excited state of the daughter nucleus is metastable, the  $\gamma$ -radiation will be emitted later.

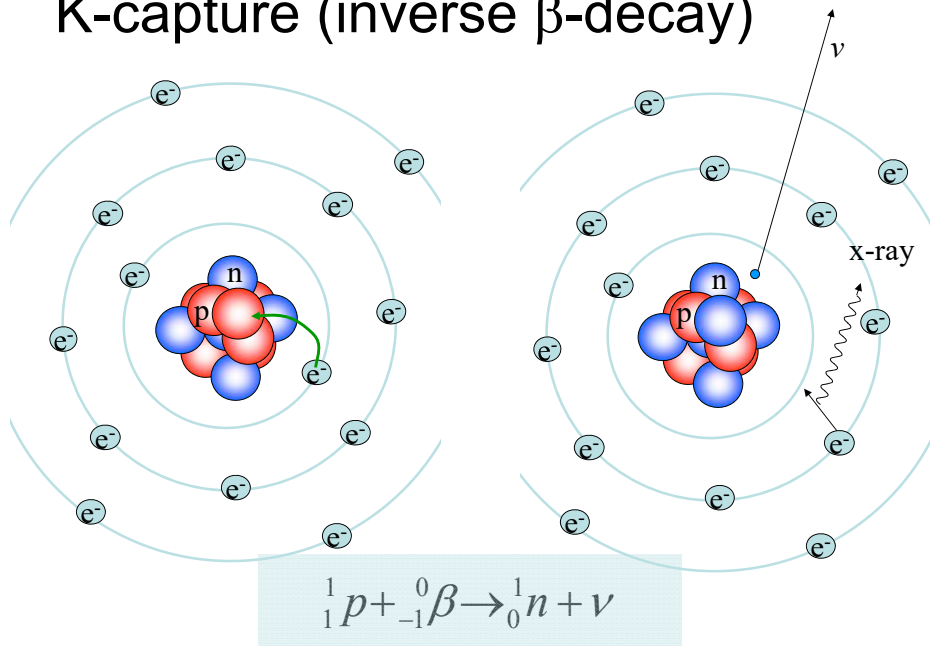
The parent and daughter atoms can be separated:  
the daughter atom emits **only  $\gamma$ -radiation!**

=> **Isotope diagnostics (nuclear imaging)**

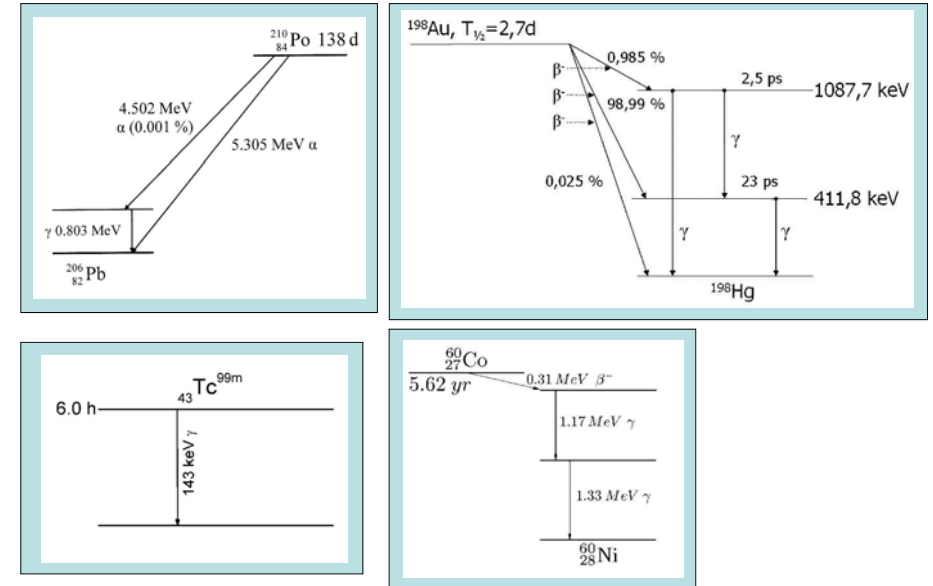
E.g.:  $^{99\text{m}}\text{Tc}$



## K-capture (inverse $\beta$ -decay)



## Some examples of the decay paths



## Characteristics of radioactive decays in general

activity	characterizes the source
half life time	characterizes the speed of the decay*
particle energy	characterizes the radiation*

\*depends on the type of the isotope

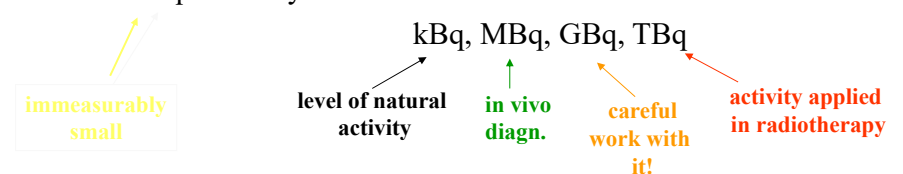
## Activity ( $\Lambda$ )

$$\Lambda = \left| \frac{dN}{dt} \right| \quad \left( = \left| \frac{\Delta N}{\Delta t} \right| \right)$$

$N$  = Number of undecayed atoms  
 $t$  = time  
 $\Delta N$  = Number of decays during  $\Delta t$  time

Activity = number of decays in a unit time

unit: becquerel Bq  
 1 Bq = 1 decay/sec



## Law of radioactive decay

$$\Delta N = -\lambda N \Delta t$$

N: Number of undecayed nuclei

$$\frac{dN}{dt} = -\lambda N$$

$\lambda$ : decay constant (probability of the decay [1/s])  
 $1/\lambda = \tau$  average lifetime

Differential equation

solution:

$$N(t) = N_0 e^{-\lambda t}$$

Exponential decrease

number of undecayed atoms at  $t=0$

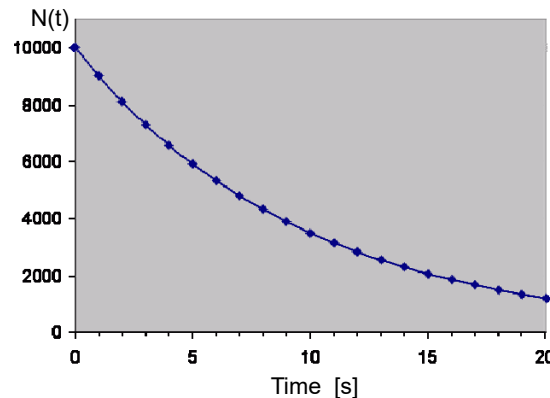
## An example

- At start:  $N_0 = 10000$  (number of atoms)  
 $\lambda = 0.1 \text{ } 1/\text{s}$  (decay probability)
- after 1 sec: 9000 ( $10000 \times 0.1 = 1000$  decayed)
- after 2 sec: 8100 ( $9000 \times 0.1 = 900$  decayed)
- after 3 sec: 7290 ( $8100 \times 0.1 = 810$  decayed)
- after 4 sec: 6561 ( $7290 \times 0.1 = 729$  decayed)
- ....

## An example

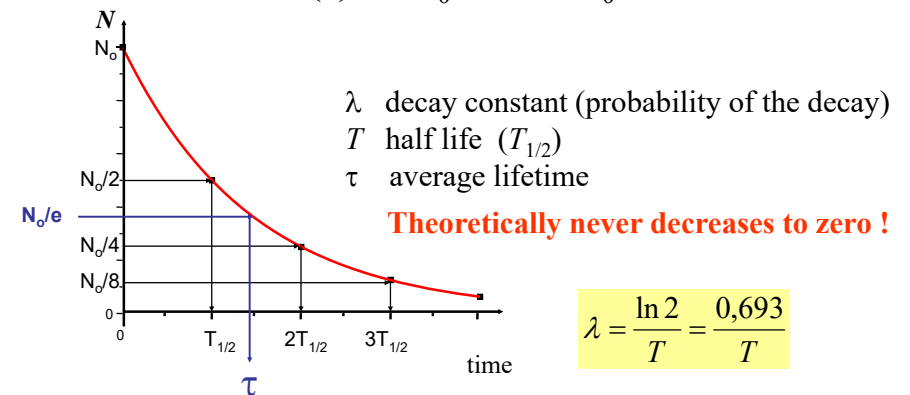
- At start:  $N_0 = 10000$  (number of atoms)  
 $\lambda = 0.1 \text{ } 1/\text{s}$  (decay probability)

- after 1 s: 9000
- after 2 s: 8100
- after 3 s: 7290
- after 4 s: 6561
- ....



## Law of radioactive decay

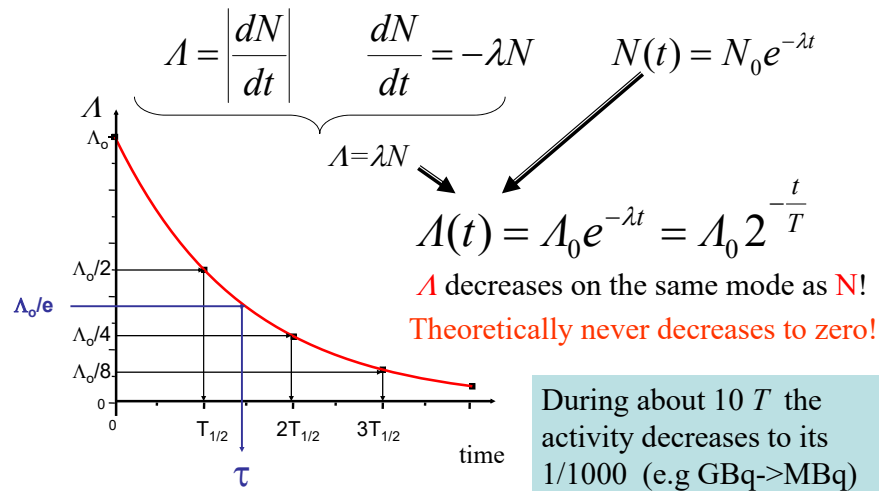
$$N(t) = N_0 e^{-\lambda t} = N_0 2^{-\frac{t}{T}}$$



$$\lambda = \frac{\ln 2}{T} = \frac{0,693}{T}$$

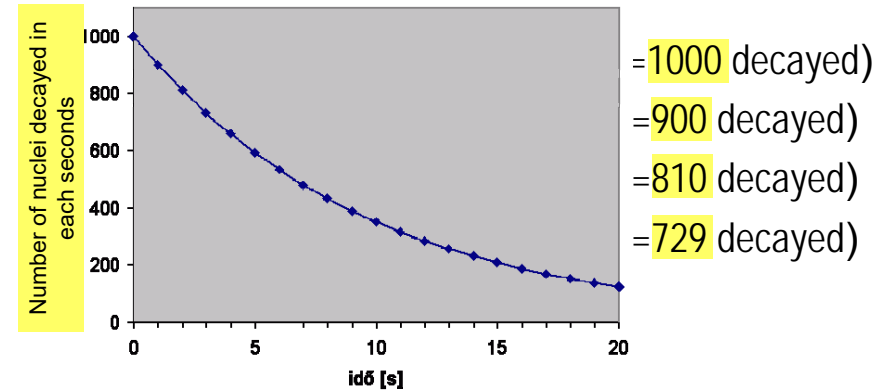


## Decrease of the activity as a function of time



## An example

- At start:  $N_0 = 10000$  (number of atoms)  
 $\lambda = 0.1 \text{ } ^1/\text{s}$  (decay probability)



## Few examples for half life

$^{232}\text{Th}$	$1.4 \cdot 10^{10} \text{ y}$	$^{60}\text{Co}$	5.3 y
$^{238}\text{U}$	$4.5 \cdot 10^9 \text{ y}$	$^{59}\text{Fe}$	1.5 m
$^{40}\text{K}$	$1.3 \cdot 10^9 \text{ y}$	$^{56}\text{Cr}$	1 m (28 d)
$^{14}\text{C}$	5736 y	$^{131}\text{I}$	8 d
$^{137}\text{Cs}$	30 y	$^{99\text{m}}\text{Tc}$	6 h
$^3\text{H}$	12.3 y	$^{18}\text{F}$	110 min
		$^{11}\text{C}$	20 min
		$^{15}\text{O}$	2 min
		$^{222}\text{Th}$	2.8 ms

**Don't learn these numbers!**

## Typical energy levels in the microworld

Excitation of the outer electrons

eV (aJ)

light



Electron transition between inner electrons

keV (fJ)

X-ray



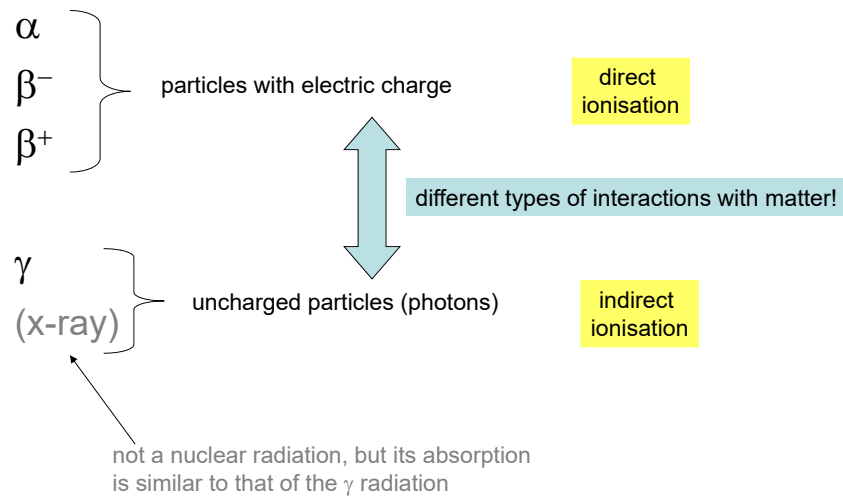
Transformation of the nucleus (decay)

MeV (pJ)

Nuclear radiation  
 $\alpha$ ,  $\beta$ ,  $\gamma$



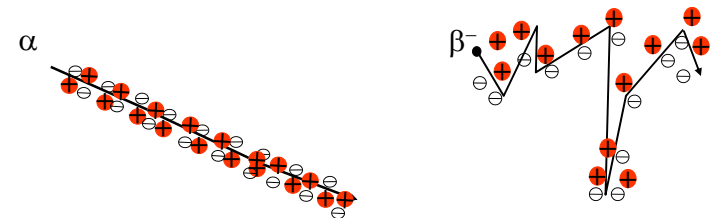
## Absorption of the nuclear radiation



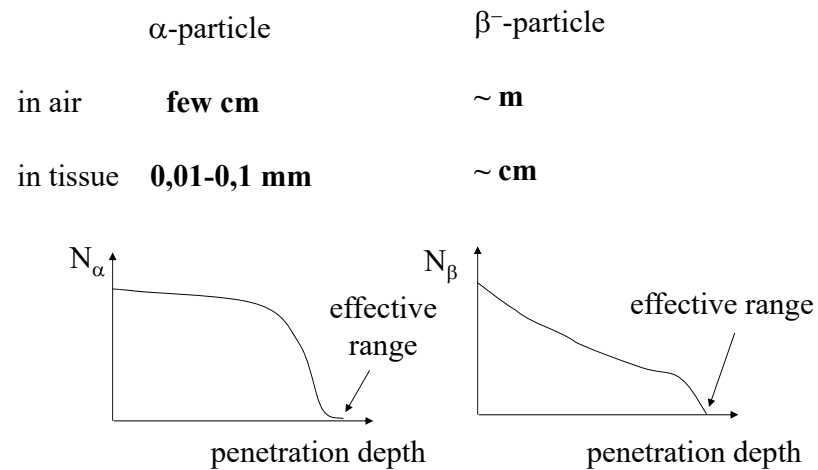
## Absorption of the charged particles

Ionizing during the path => continuous decrease of the particle energy  
The energy after a given path length decreases to the thermal value

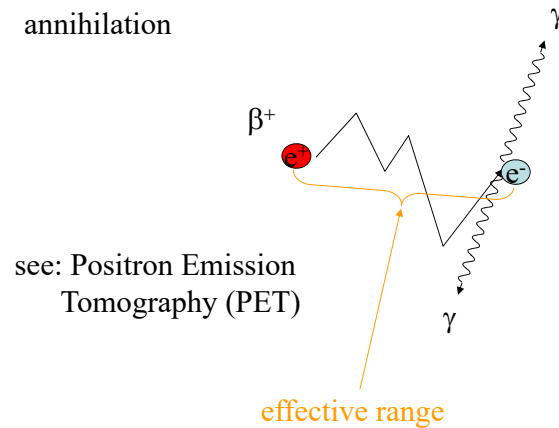
effective range



## Effective range



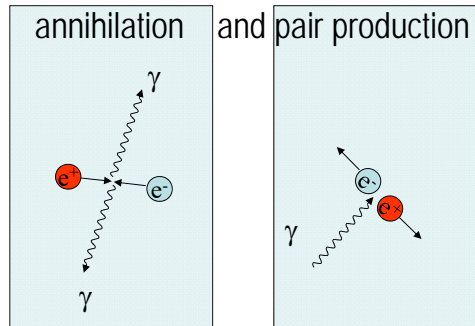
## $\beta^+$ -radiation





# Electron and positron

- particle - antiparticle
- same mass,
- charge: same value, but different

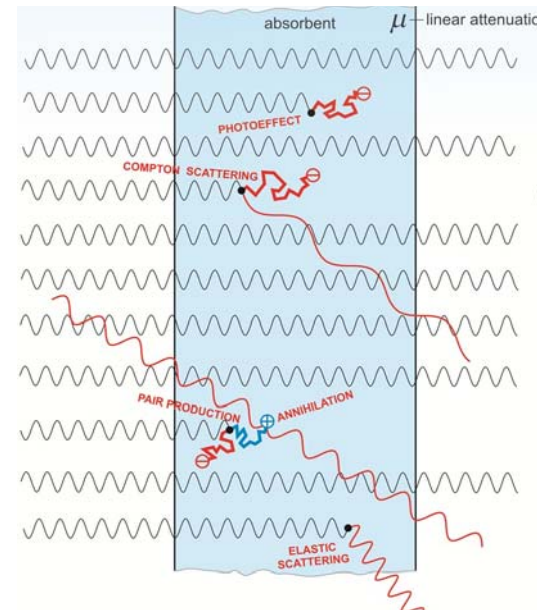


Einstein:  
mass-energy  
equivalence

$$E=mc^2$$

$$m_e c^2 = 511 \text{ keV} \approx 0.5 \text{ MeV}$$

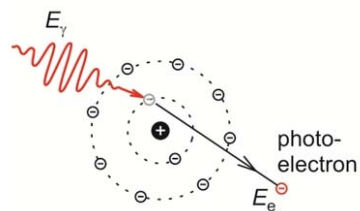
## Absorption of the $\gamma$ -radiation (and x-ray)



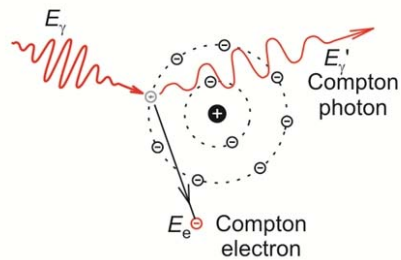
Absorption processes  
happen accidentally :

Photoeffect,  
Compton-effect,  
Pair production,  
(elastic scattering)

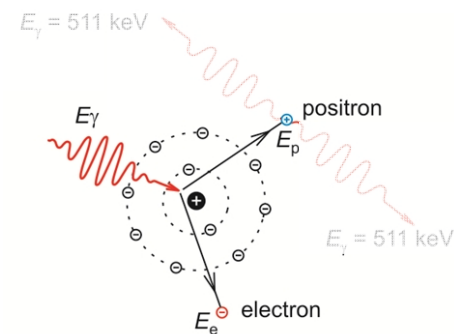
### Photoeffect



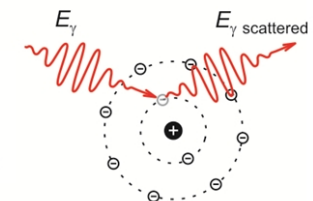
### Compton effect Compton-scattering



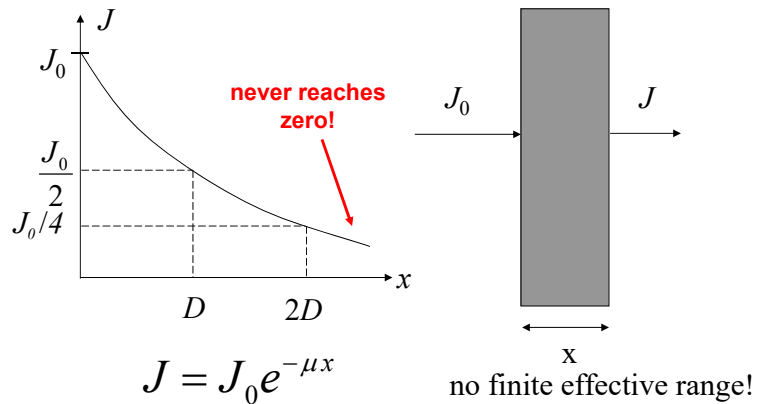
### Pair production



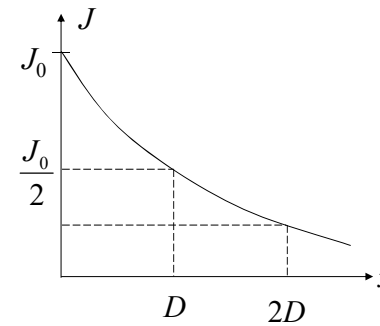
### elastic scattering



# Attenuation of the $\gamma$ -radiation and x-ray



few rules of thumb:  $x_{1/10} = 3,33 D$      $x_{1/1000} = 10 D$



$$J = J_0 e^{-\mu x}$$

$\mu$ : (linear) attenuation coefficient

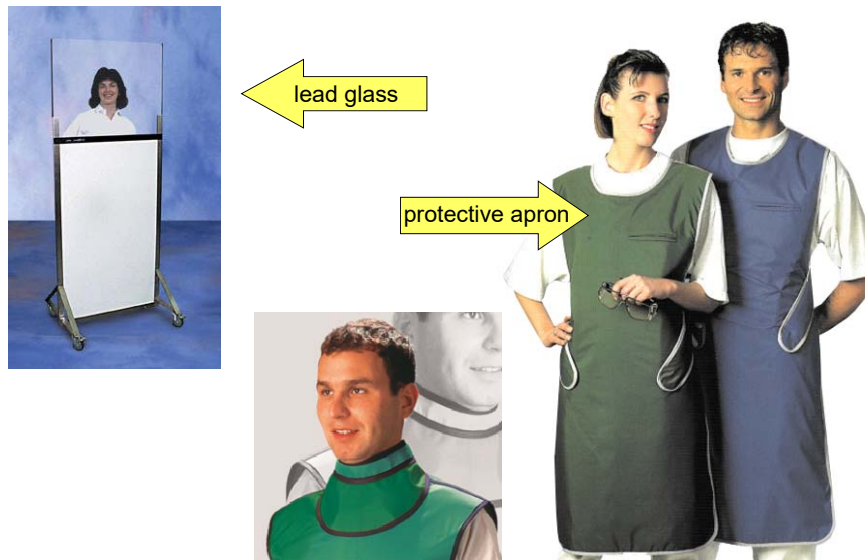
its units are: 1/m, 1/cm

$\delta = \frac{1}{\mu}$  „penetration depth”  
Intensity decreases to the e-th part (c.a. 37%)

$\mu$ (material, number of absorbing centers, energy of the radiation)  
 $= \mu(\text{material}, \rho, E_{\text{photon}}) \sim \rho$

$$\mu_m = \frac{\mu}{\rho} \quad \text{mass attenuation coefficient}$$

## Applications (attenuation)



## Applications: isotopes and nuclear radiation

