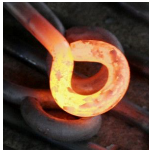


# Generation of light

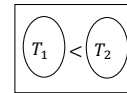
## Thermal radiation and Luminescence



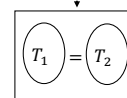
Zsolt Mártonfalvi

1

## Thermal radiation



The temperature of the two bodies will become equal with time even in vacuum!  
(no convection or conduction between them)



Consequently:  
All bodies emit radiation independently of the temperature of their surroundings. The emitted radiation is always **electromagnetic radiation**.

Quantitative description of thermal radiation:

- **Radiant emittance ( $M$ )**

$$M = \frac{\Delta P}{\Delta A} = \left[ \frac{W}{m^2} \right]$$

- **Absorption coefficient ( $\alpha$ )**

$$\alpha = \frac{I_{\text{absorbed by the surface}}}{I_{\text{received by the surface}}} \quad (0 \leq \alpha \leq 1)$$

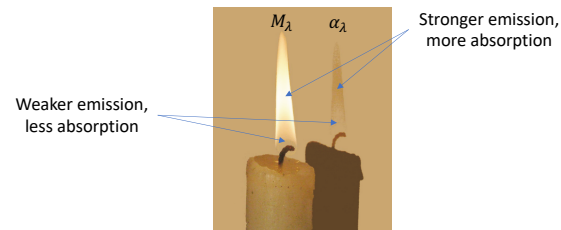
$M$  and  $\alpha$  strongly depend on the absolute temperature of the body!

2

## Kirchhoff's law

Bodies that emit more also absorb more. The ratio between radiant emittance and absorption coefficient is constant within a narrow range of wavelength ( $\lambda$ ):

$$\frac{M_{\lambda \text{ body1}}}{\alpha_{\lambda \text{ body1}}} = \frac{M_{\lambda \text{ body2}}}{\alpha_{\lambda \text{ body2}}} = \text{constant}$$



3

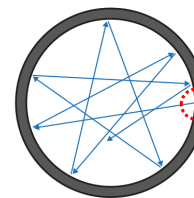
## The ideal black body

A theoretical body that is used as a model in the description of thermal radiation. It absorbs all radiation that falls on it:

$$\alpha_{\text{black body}} = 1$$

thus we can calculate any real body's radiant emittance if we know its absorption coefficient ( $\alpha_{\lambda i}$ ):

$$M_{\lambda i} = \alpha_{\lambda i} M_{\text{black body}}$$



The hole on a dark cavity approximates an ideal black body

The radiant emittance of a black body depends strongly on the absolute temperature!

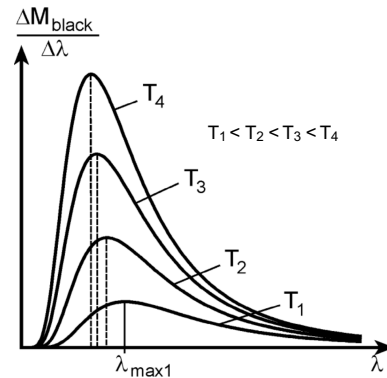
**Stefan's law:**

$$M = \sigma T^4 \quad \sigma = 5.67 \cdot 10^{-8} \frac{W}{m^2 K^4}$$

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## The ideal black body

The emission spectrum of the black body is **continuous**.



Wien's displacement law:

$$\lambda_{max} T = \text{constant}$$

Max Planck's radiation law:

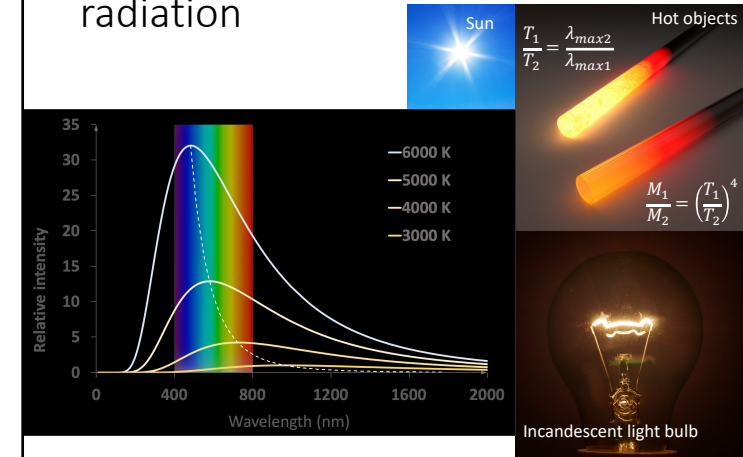
The radiation energy of a black body is emitted in discrete integer multiples of a fundamental "package", the quantum.

$$E = hf$$

$$h = 6.62 \cdot 10^{-34} \text{ Js}$$

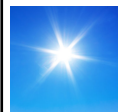
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## Light sources based on thermal radiation

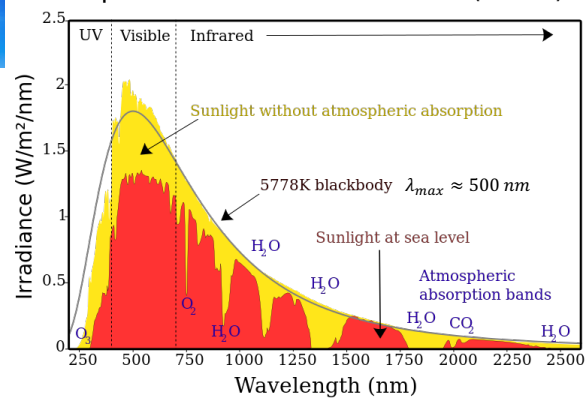


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## Sun light



### Spectrum of Solar Radiation (Earth)



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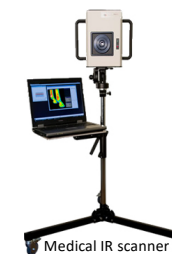
## Medical applications

### Telethermography

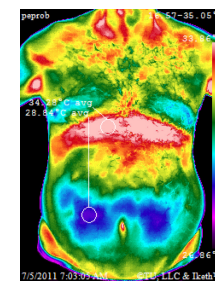
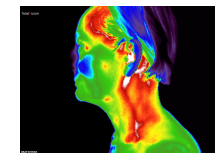
Measures the radiant emittance of thermal radiation of the human body.

Human body:  $\lambda_{max} \approx 10 \mu\text{m}$  (infrared)  
 $\alpha \approx 0.95$  (95% of ideal black body)

Radiant emittance strongly depends on temperature!



Diagnosis of inflammations, cancer, circulatory defects that cause the change in local temperature.

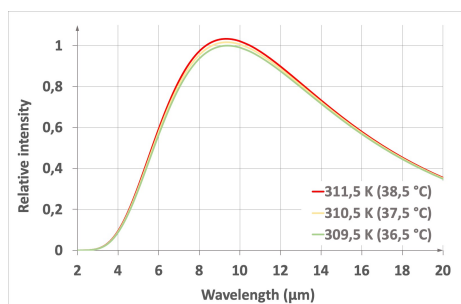


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## Medical applications

### Non-contact thermometry

1 °C of temperature difference will increase the intensity by ~1.5% at the emission maximum!



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## Luminescence

Spontaneous emission of a photon due to the relaxation of an excited electron.

Phases of luminescence:

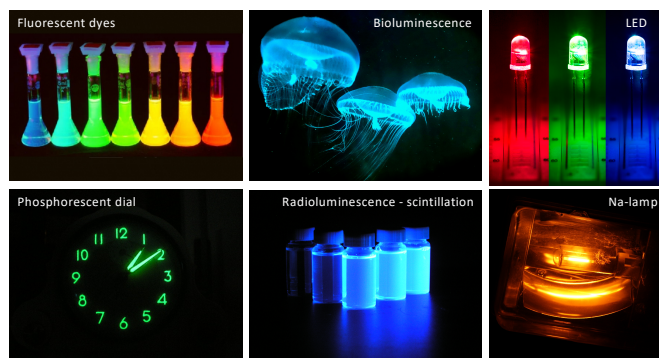
- Absorption of external energy
- Excitation
- Emission of energy in the form of electromagnetic radiation

### Types of luminescence

Type of excitation	Name	Example
Light	Photoluminescence	Fluorescent lamp
Electric	Electroluminescence	Hg vapor lamp
Radioactive	Radioluminescence	Nal (Tl)
Mechanical	Triboluminescence	(Sugar cube)
Biochemical	Bioluminescence	firefly
Thermal	Thermoluminescence	CaSO <sub>4</sub> (Dy) (dosimeter)

Type of relaxation	Name	Example
S <sub>1</sub> – S <sub>0</sub> (fast)	Fluorescence	Fluorescein
T <sub>1</sub> – S <sub>0</sub> (slow)	Phosphorescence	Phosphorous

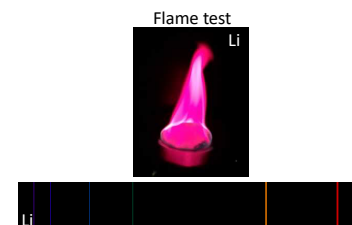
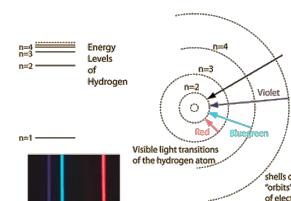
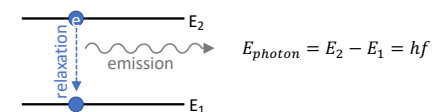
10



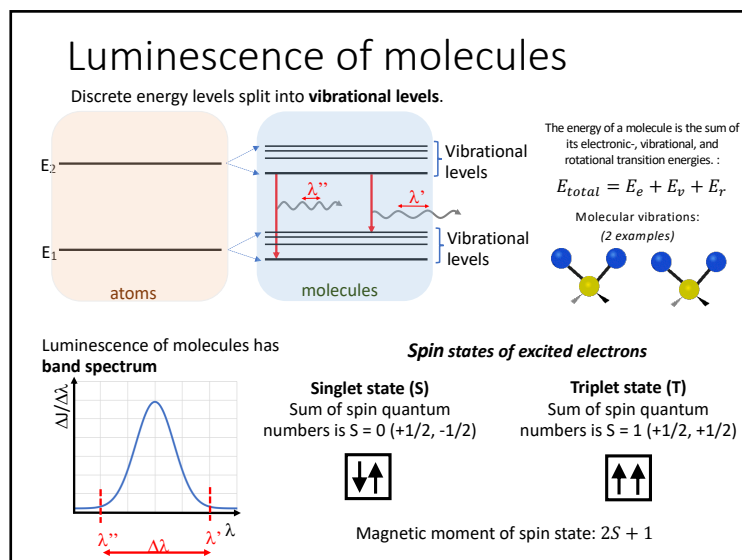
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## Luminescence of atoms

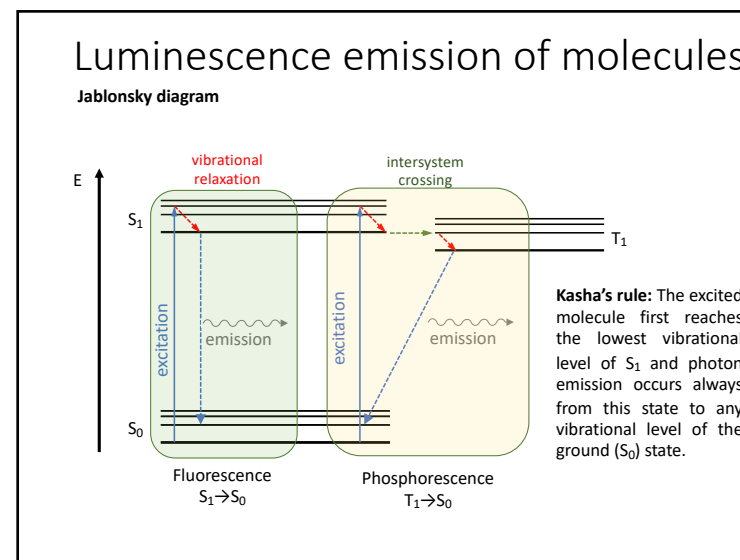
Luminescence emission of atoms has a **line spectrum**.



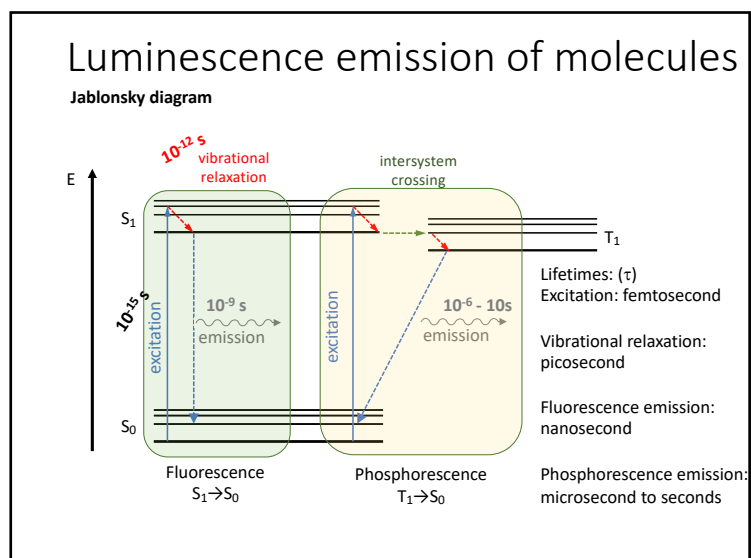
12



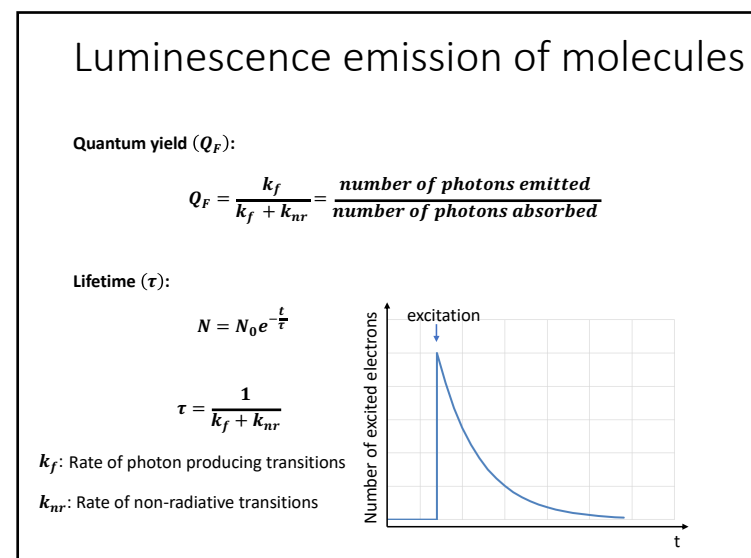
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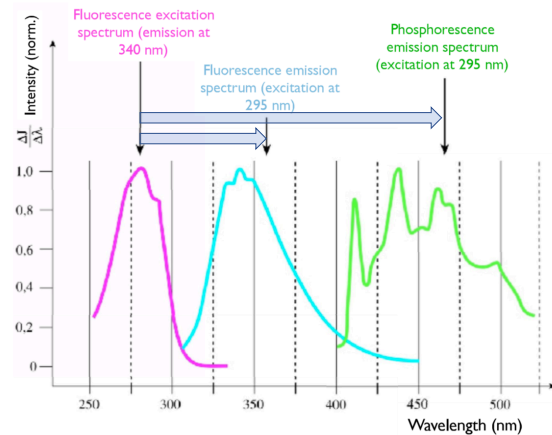


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## Stokes shift

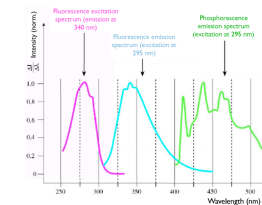
The shift between the emission and excitation spectra



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## Comparing emission processes

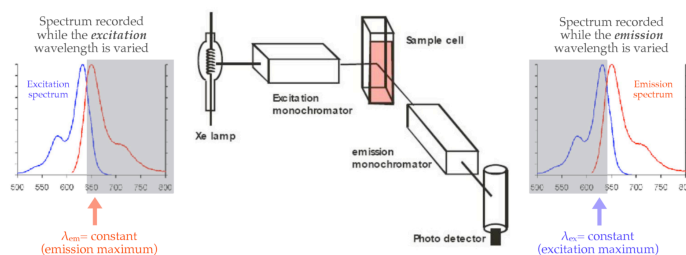
	Fluorescence	Phosphorescence
Relaxation	From singlet state $S_1 \rightarrow S_0$	From triplet state $T_1 \rightarrow S_0$
Lifetime	Nanoseconds	From microseconds up to seconds
Stokes shift	Smaller	Larger (because $T_1$ is at lower energy than $S_1$ )



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## Measurement of luminescence

Luminescence spectrofluorimeter

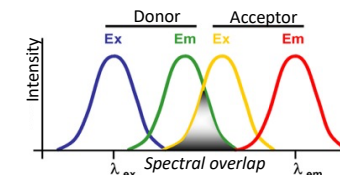


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## Applications - FRET

Förster Resonance Energy Transfer

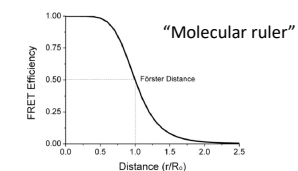
Energy transfers from donor without emission to acceptor in dipole-dipole interactions. Requires spectral overlap between donor emission and acceptor absorption.



FRET efficiency or quantum yield ( $E$ ):

$$E = \frac{1}{1 + (r/r_0)^6}$$

$r_0$ : Förster distance

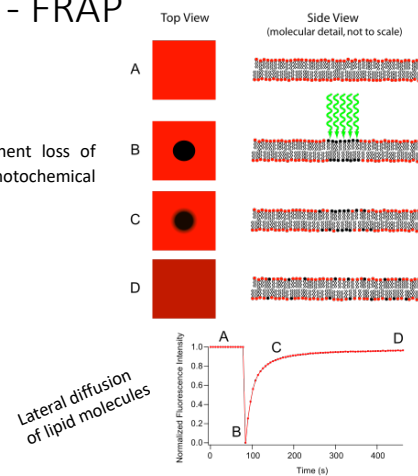


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## Applications - FRAP

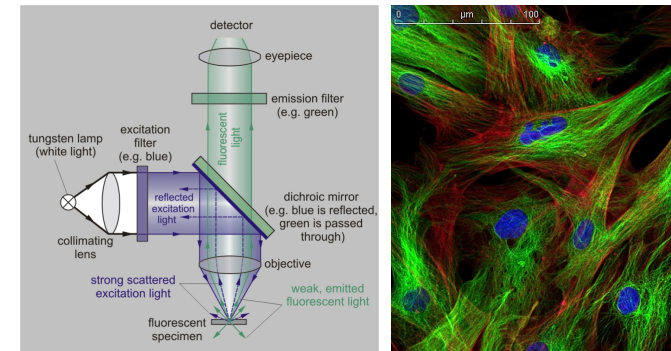
### Fluorescence Recovery After Photobleaching

Photobleaching: The permanent loss of fluorescence due to photochemical reactions.



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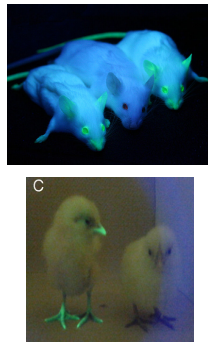
## Fluorescence microscope



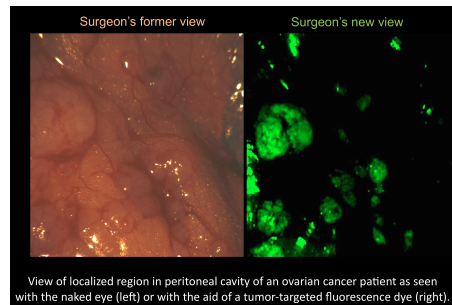
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## Biomedical applications

### Fluorescent transgenic animals



### Fluorescence guided surgery



View of localized region in peritoneal cavity of an ovarian cancer patient as seen with the naked eye (left) or with the aid of a tumor-targeted fluorescence dye (right).

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