

# Generation of light

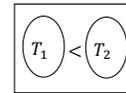
*Thermal radiation and Luminescence*



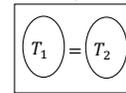
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1

# Thermal radiation



The temperature of the two bodies will become equal with time even in vacuum! (no convection or conduction between them)



Consequently: All bodies emit radiation independently of the temperature of their surroundings. The emitted radiation is always **electromagnetic radiation**.

Quantitative description of thermal radiation:

- **Radiant emittance ( $M$ )**

$$M = \frac{\Delta P}{\Delta A} = \left[ \frac{W}{m^2} \right]$$

- **Absorption coefficient ( $\alpha$ )**

$$\alpha = \frac{J_{\text{absorbed by the surface}}}{J_{\text{received by the surface}}} \quad (0 \leq \alpha \leq 1)$$

$M$  and  $\alpha$  strongly depend on the absolute temperature of the body!

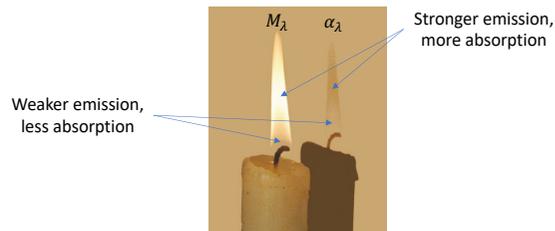
2

# Kirchhoff's law

Bodies that emit more also absorb more. The ratio between radiant emittance and absorption coefficient is constant within a narrow range of wavelength ( $\lambda$ ):



$$\frac{M_{\lambda \text{ body1}}}{\alpha_{\lambda \text{ body1}}} = \frac{M_{\lambda \text{ body2}}}{\alpha_{\lambda \text{ body2}}} = \text{constant}$$



3

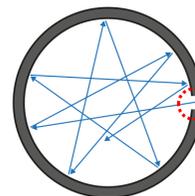
# The ideal black body

A theoretical body that is used as a model in the description of thermal radiation. It absorbs all radiation that falls on it:

$$\alpha_{\text{black body}} = 1$$

thus we can calculate any real body's radiant emittance if we know its absorption coefficient ( $\alpha_{\lambda i}$ ):

$$M_{\lambda i} = \alpha_{\lambda i} M_{\lambda \text{ black body}}$$



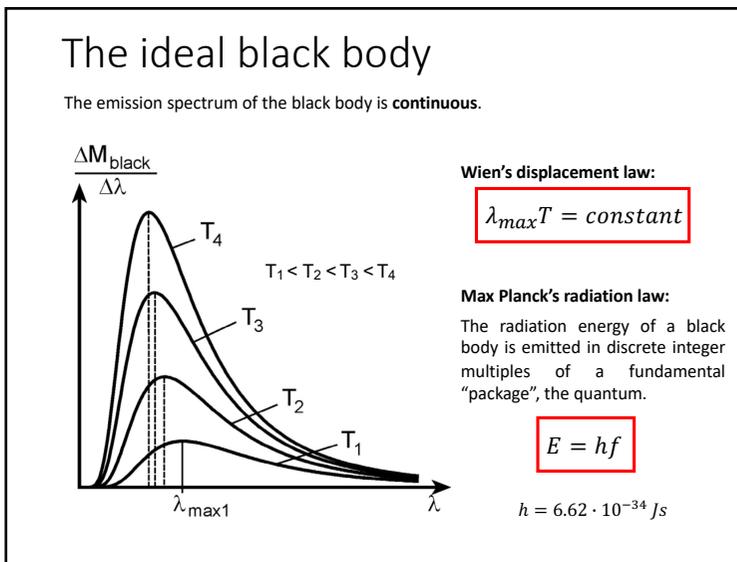
The hole on a dark cavity approximates an ideal black body

The radiant emittance of a black body depends strongly on the absolute temperature!

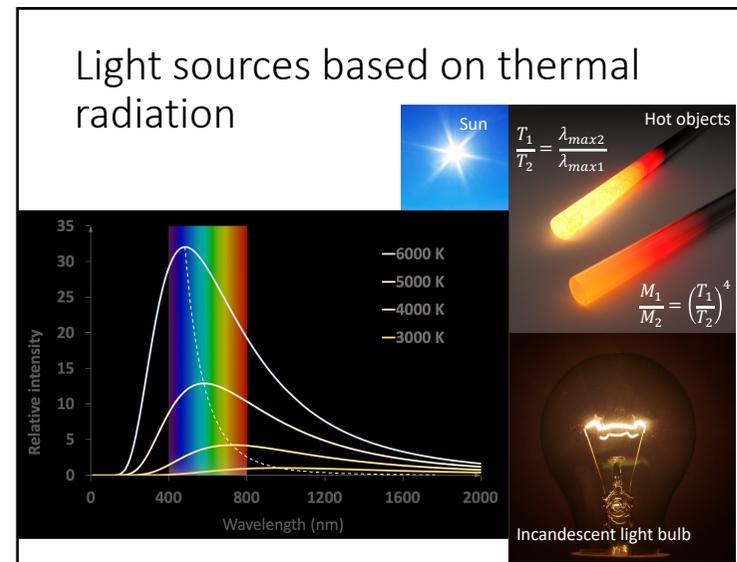
**Stefan's law:**

$$M = \sigma T^4 \quad \sigma = 5.67 \cdot 10^{-8} \frac{W}{m^2 K^4}$$

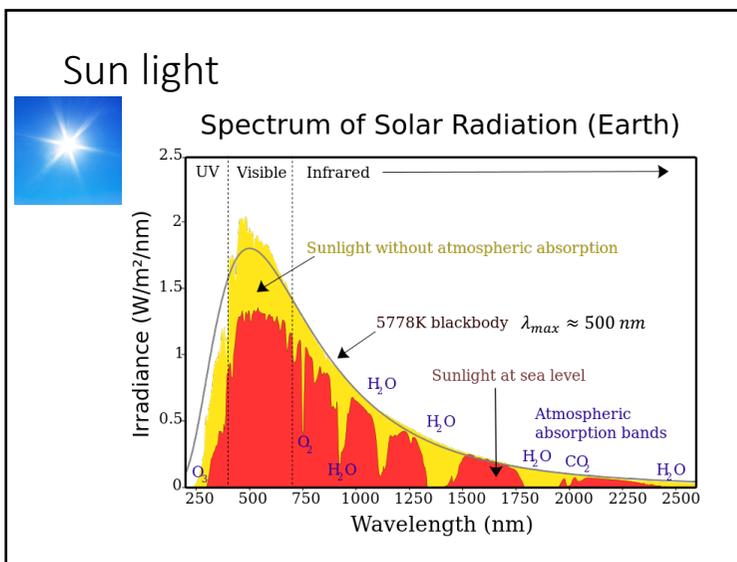
4



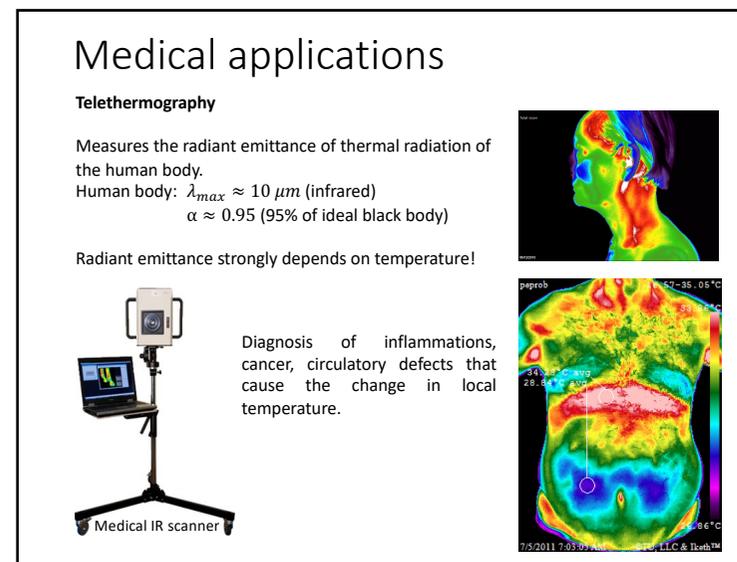
5



6



7



8

## Medical applications

**Non-contact thermometry**

1 °C of temperature difference will increase the intensity by ~1.5% at the emission maximum!

9

## Luminescence

Spontaneous emission of a photon due to the relaxation of an excited electron.

Phases of luminescence:

- Absorption of external energy
- Excitation
- Emission of energy in the form of electromagnetic radiation

**Types of luminescence**

Type of excitation	Name	Example
Light	Photoluminescence	Fluorescent lamp
Electric	Electroluminescence	Na-lamp
Radioactive	Radioluminescence	Nal (Tl) (scintillator)
Mechanical	Triboluminescence	(Percussion, friction)
Biochemical	Bioluminescence	firefly
Thermal	Thermoluminescence	CaSO <sub>4</sub> (Dy) (dosimeter)

Type of relaxation	Name	Example
S <sub>1</sub> – S <sub>0</sub> (fast)	Fluorescence	Fluorescein
T <sub>1</sub> – S <sub>0</sub> (slow)	Phosphorescence	Phosphorous

10

11

## Luminescence of atoms

Luminescence emission of atoms has a **line spectrum**.

12

### Luminescence of molecules

Discrete energy levels split into **vibrational levels**.

The energy of a molecule is the sum of its electronic, vibrational, and rotational transition energies:  $E_{total} = E_e + E_v + E_r$

Molecular vibrations: (2 examples)

Luminescence of molecules has **band spectrum**

**Spin states of excited electrons**

**Singlet state (S)**  
Sum of spin quantum numbers is  $S = 0$  (+1/2, -1/2)

**Triplet state (T)**  
Sum of spin quantum numbers is  $S = 1$  (+1/2, +1/2)

Magnetic moment of spin state:  $2S + 1$

13

### Luminescence emission of molecules

**Jablonsky diagram**

**Kasha's rule:** The excited molecule first reaches the lowest vibrational level of  $S_1$  and photon emission occurs always from this state to any vibrational level of the ground ( $S_0$ ) state.

14

### Luminescence emission of molecules

**Jablonsky diagram**

**Lifetimes ( $\tau$ ):**  
Excitation: femtosecond  
Vibrational relaxation: picosecond  
Fluorescence emission: nanosecond  
Phosphorescence emission: microsecond to seconds

15

### Luminescence emission of molecules

**Quantum yield ( $Q_F$ ):**

$$Q_F = \frac{k_f}{k_f + k_{nr}} = \frac{\text{number of photons emitted}}{\text{number of photons absorbed}}$$

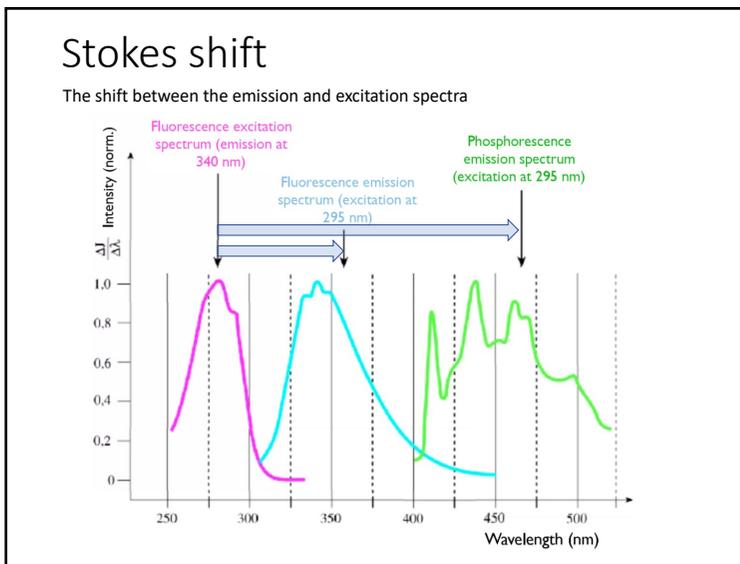
**Lifetime ( $\tau$ ):**

$$N = N_0 e^{-\frac{t}{\tau}}$$

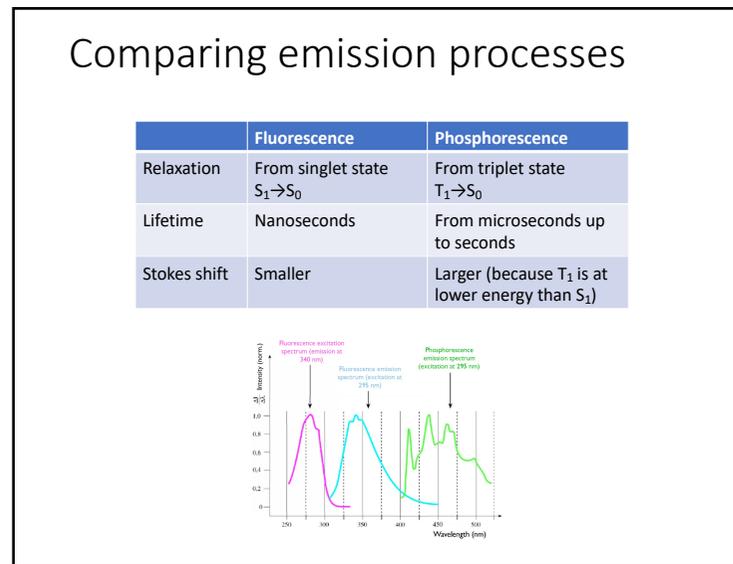
$$\tau = \frac{1}{k_f + k_{nr}}$$

$k_f$ : Rate of photon producing transitions  
 $k_{nr}$ : Rate of non-radiative transitions

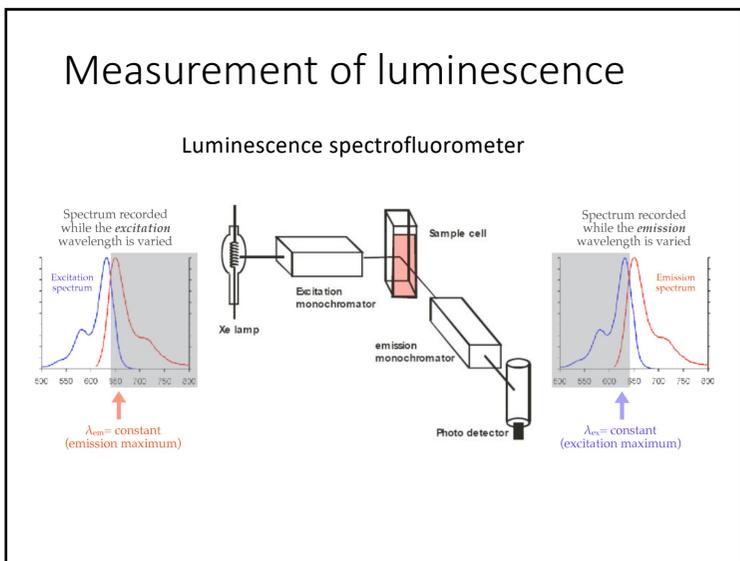
16



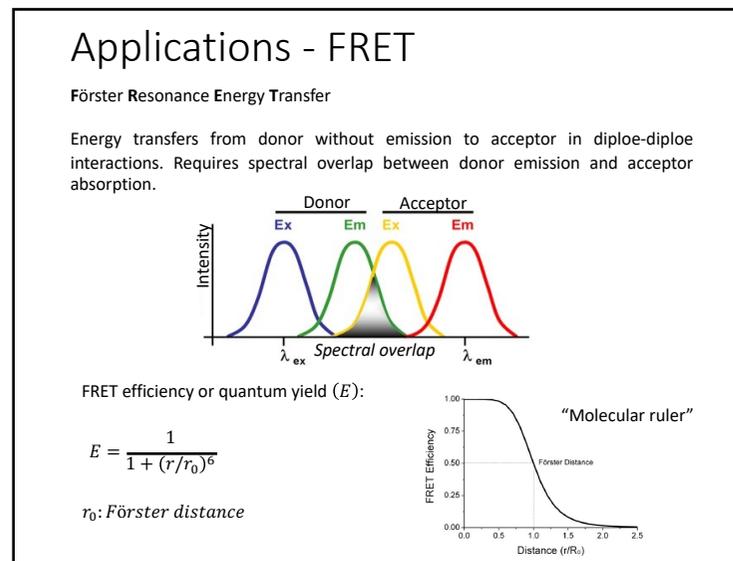
17



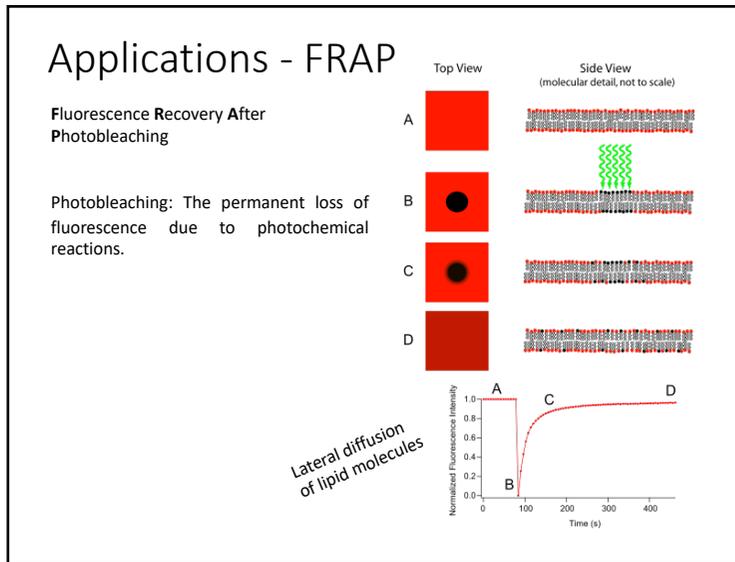
18



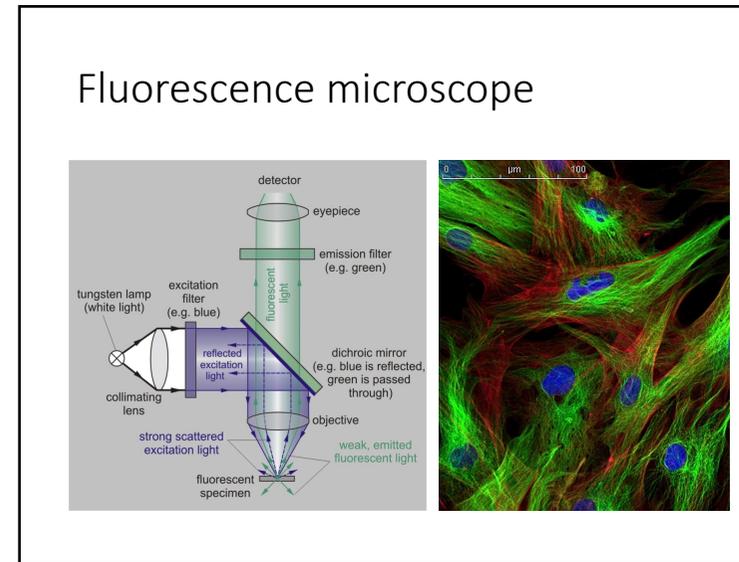
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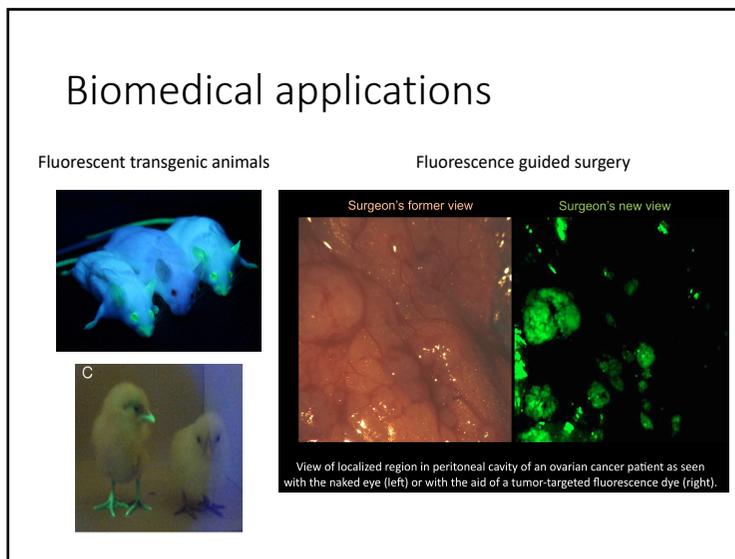
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21



22



23