



Physical Foundations of Dental Materials Science

Introduction

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Important informations

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Head: Prof. Miklós Kellermayer
- <http://itc.semmelweis.hu/>
- Pdf format e -book (Physical bases of dental material science)
- Exam: written test composed of three sections:
1) Definitions, 2) Calculations, 3) Theory

Further readings:

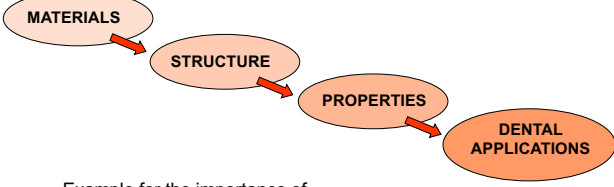
- W.D. Callister: *Materials Science and Engineering. An Introduction* (7th ed.), Wiley&Sons, 2007
- K.J. Anusavice: *Phillips' Science of Dental Materials* (11th ed.), Saunders, 2003
- Damjanovich, Fidy, Szöllösi: *Medical Biophysics*, Medicina, Budapest, 2009

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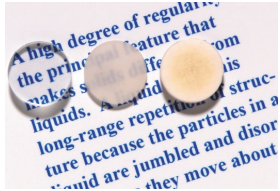
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How to start? – How to proceed?

The way how the lectures proceed




Example for the importance of structure:



All are Al_2O_3 !

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Physical Foundations of Dental Materials Science

1.

Structure of matter

Atomic interaction, multiatomic system - gases

E-book chapters:

1, 2, 3

Highlights:

- ❖ Interactions
- ❖ Energy curve of atomic and molecular interactions
- ❖ Interpretation of temperature
- ❖ Boltzmann-distribution




Problems:

Chapter 1, 2, 3.:
1, 3, 9, 10, 13, 17, 19


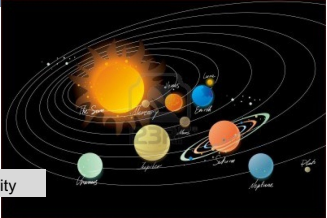
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Interactions, their role and description






in the background of contacts:
molecular interactions

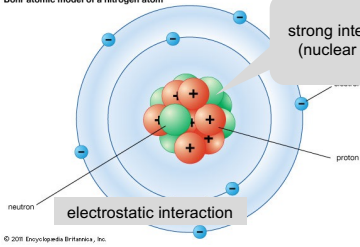



gravity

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Bohr atomic model of a nitrogen atom



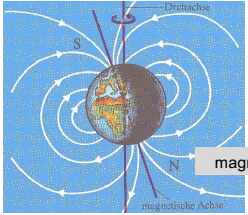
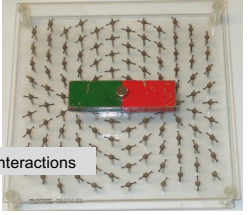
strong interactions (nuclear forces)

proton

neutron

electrostatic interaction

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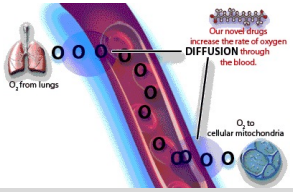
magnetic interactions

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thermal interaction (heat)



Our novel drugs increase the rate of oxygen DIFFUSION through the blood.

O₂ from lungs

O₂ to cellular mitochondria

chemical, biological interactions

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How bodies are formed in general:

INTERACTIONS

REPULSIVE

MOVEMENT

ATTRACTIVE

„disorder“

→

←

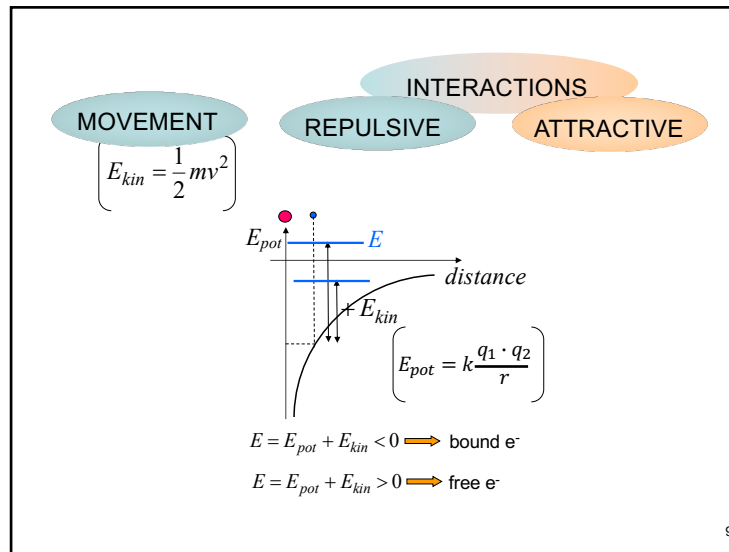
„order“

How to compare these quantitatively?

ENERGY

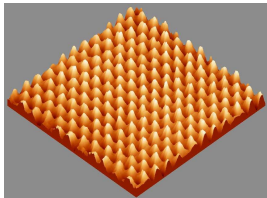
Energy of an interaction (potential energy), kinetic energy

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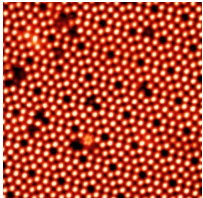


All matter is composed of atoms

- Democritus B.C. 5th century
- Dalton's atomic theory 1803
- Rutherford 1911
- Bohr 1913

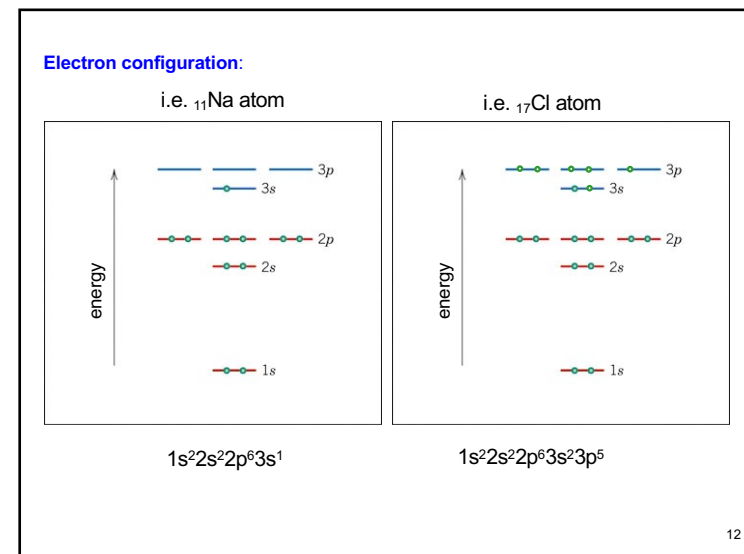
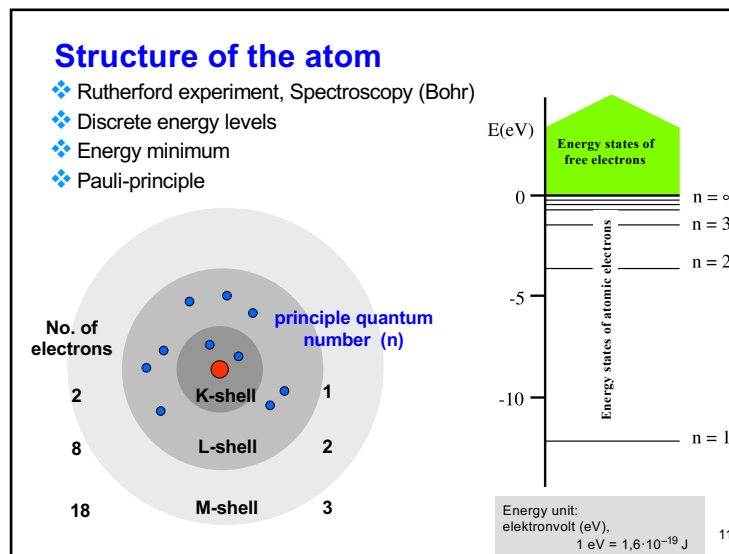


C atoms in crystal lattice
no vacancies



Si crystal with
vacancies

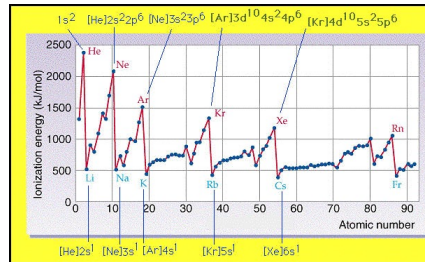
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Electronegativity

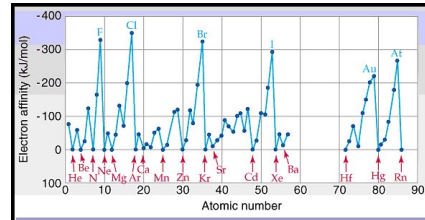
Ionization energy (I):

The amount of energy that is necessary to remove the most loosely bound electron from an atom (eV/atom; kJ/mol)



Electronaffinity (A):

The amount of energy released when an electron is added to an atom (eV/atom; kJ/mol)

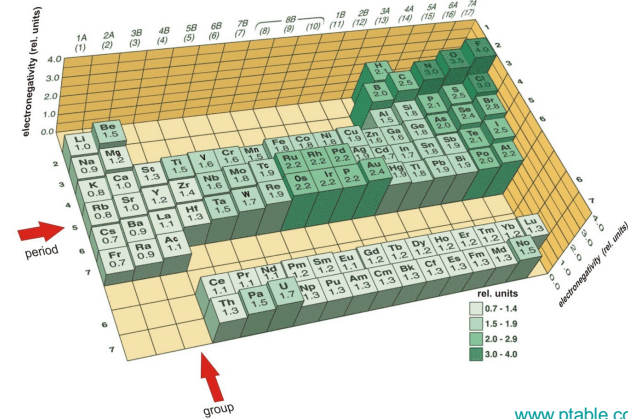


Electronegativity (EN):

$$EN = I + |A|$$

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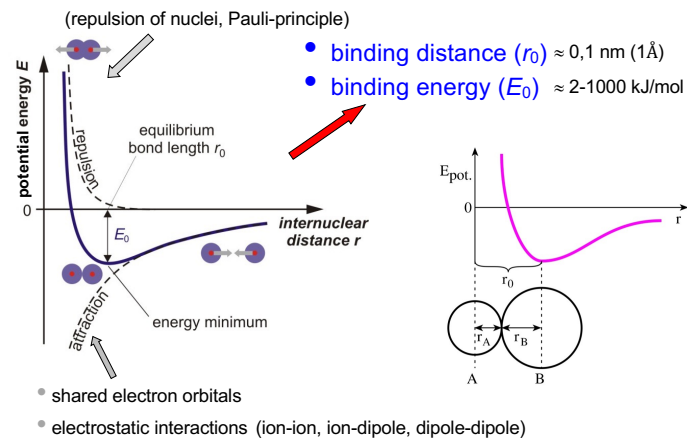
Pauling-scale:



www.ptable.com

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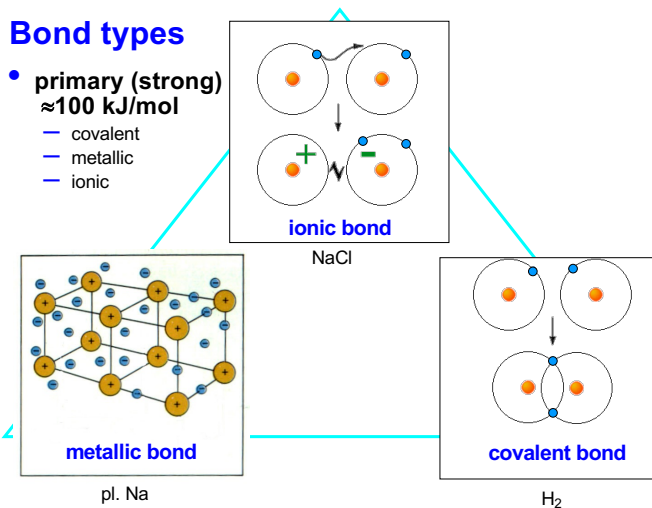
Atomic interactions



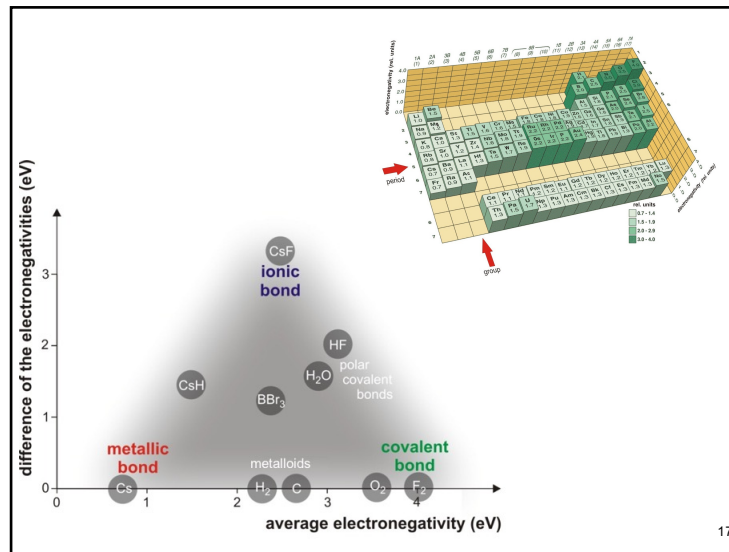
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Bond types

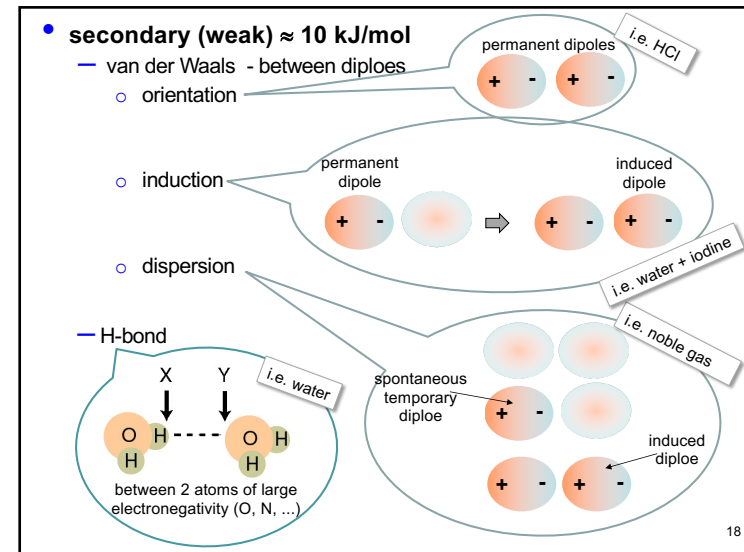
- **primary (strong)**
 ≈ 100 kJ/mol
- covalent
- metallic
- ionic



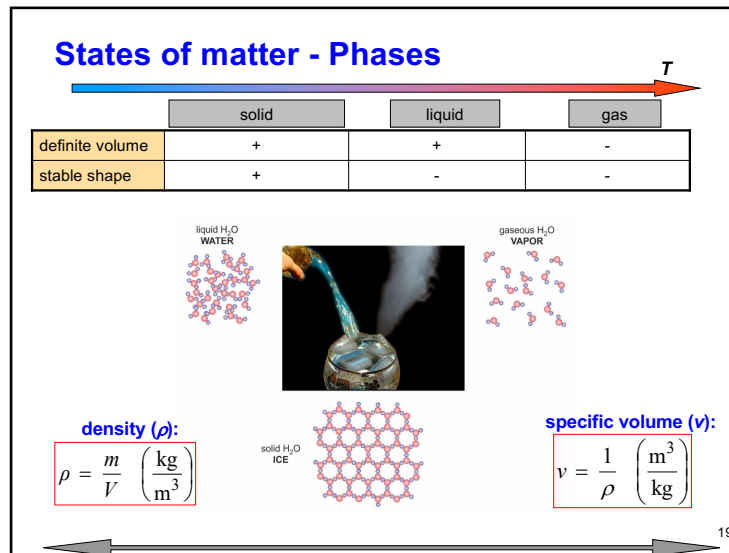
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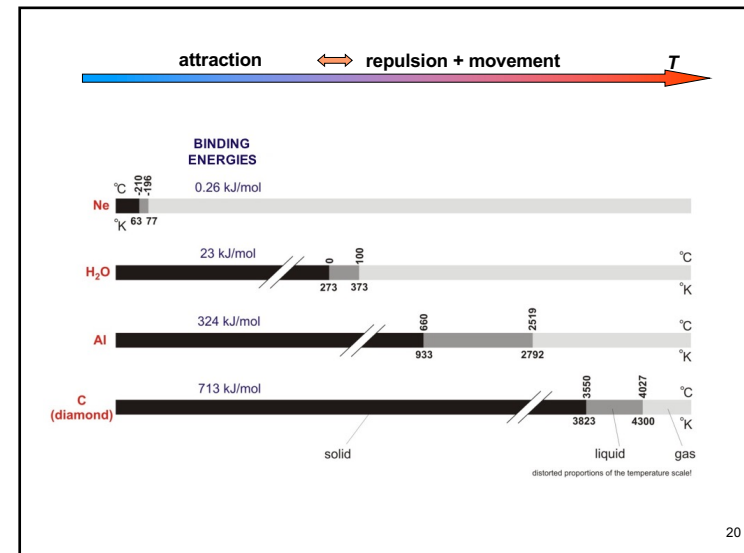
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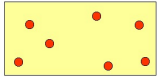


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Gases



Macroscopic description:

- No definite volume or shape
- isotropic

$$p, V, n, T$$

$$pV = nRT$$

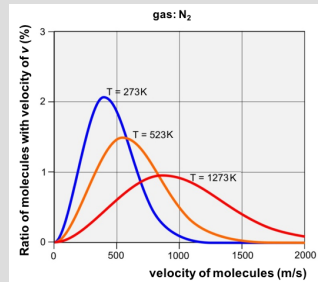
(ideal gas)

Microscopic description:

- random
- movement in many degrees of freedom

$$\frac{1}{2} m \overline{v^2} = \frac{3}{2} kT$$

Maxwell-Boltzmann- distribution



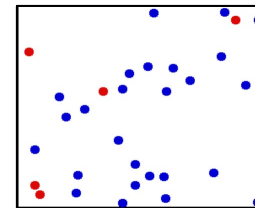
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Temperature

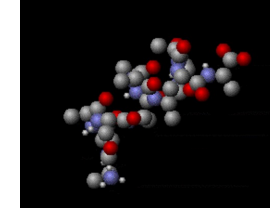
Temperature is a parameter proportional to the average kinetic energy available for each degree of freedom.

$$T(K) = t(^{\circ}C) + 273$$

hot and cold gas particles



macromolecule



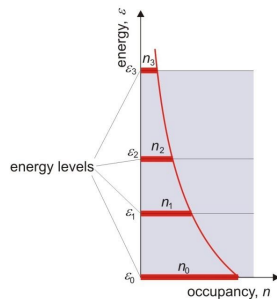
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Boltzmann-distribution

The Boltzmann distribution describes the distribution of the particles between energy levels in a force field in case of thermal equilibrium.

$$\left. \begin{array}{l} n_i \\ n_0 \end{array} \right\} \begin{array}{l} \varepsilon_i \\ \varepsilon_0 \end{array} \Delta \varepsilon$$

$$n_i = n_0 \cdot e^{-\frac{\varepsilon_i - \varepsilon_0}{kT}}$$



$$\left(\begin{array}{l} n_i = n_0 \cdot e^{-\frac{\Delta \varepsilon}{kT}} = n_0 \cdot e^{-\frac{\Delta E}{RT}} \\ \Delta E = \Delta \varepsilon \cdot N_A \\ R = k \cdot N_A \end{array} \right)$$

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Gas in a force field – gravitation

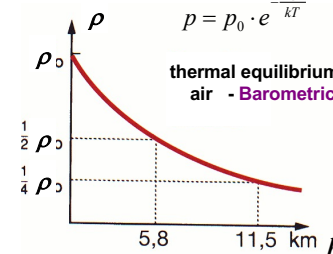
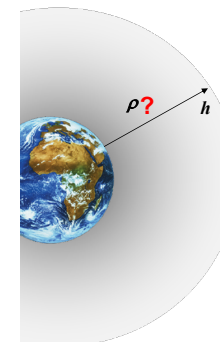
Example: density (ρ) of air **changes** in function of the potential energy

$$\rho = \rho_0 \cdot e^{-\frac{mgh}{kT}}$$

$$p = \text{const} \cdot \rho$$

$$p = p_0 \cdot e^{-\frac{mgh}{kT}}$$

thermal equilibrium !
air - **Barometric Formula**



less and less particles are found (in the same volume) at higher potential energy (at higher altitudes)

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Examples for Boltzmann-distribution:

- barometric formula
- thermal emission of electrons from metals
- Nernst-equation
- rate of chemical reactions
- conductivity of semiconductors
- number of vacancies in a metal
- ...

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