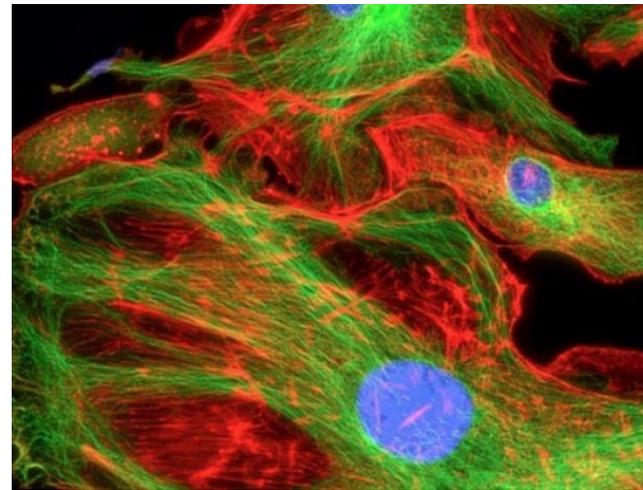


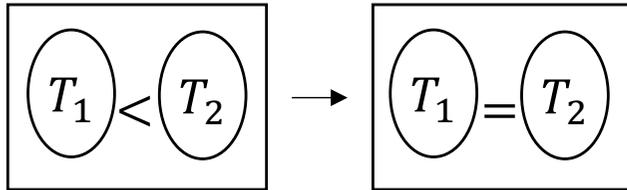
# Mechanisms of light emission

*Laws of thermal radiation and Luminescence*



Zsolt Mártonfalvi

# Thermal radiation

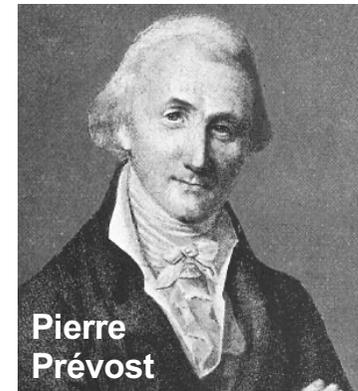


The temperature of the two bodies will become equal with time even in vacuum!  
(no convection or conduction between them)

Consequently:

All bodies emit radiation independently of the temperature of their surroundings. (*Prévost, 1791*)

The emitted radiation is always **electromagnetic radiation**.



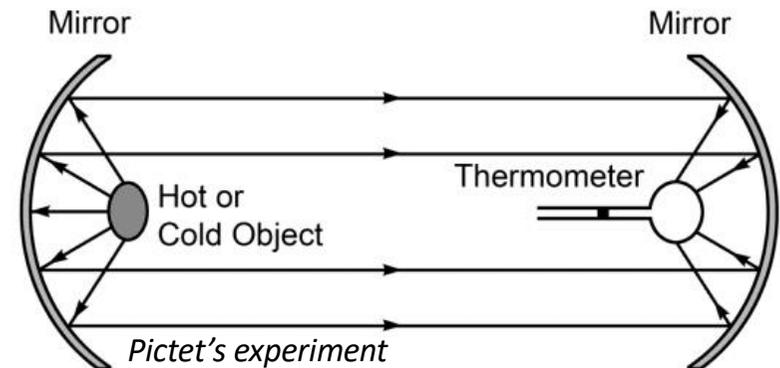
Quantitative description of thermal radiation:

- **Radiant emittance ( $M$ )**

$$M = \frac{\Delta P}{\Delta A} = \left[ \frac{W}{m^2} \right]$$

- **Absorption coefficient ( $\alpha$ )**

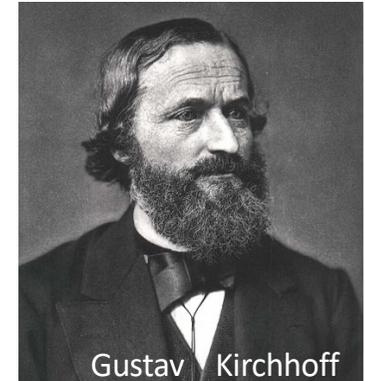
$$\alpha = \frac{J_{\text{absorbed by the surface}}}{J_{\text{received by the surface}}} \quad (0 \leq \alpha \leq 1)$$



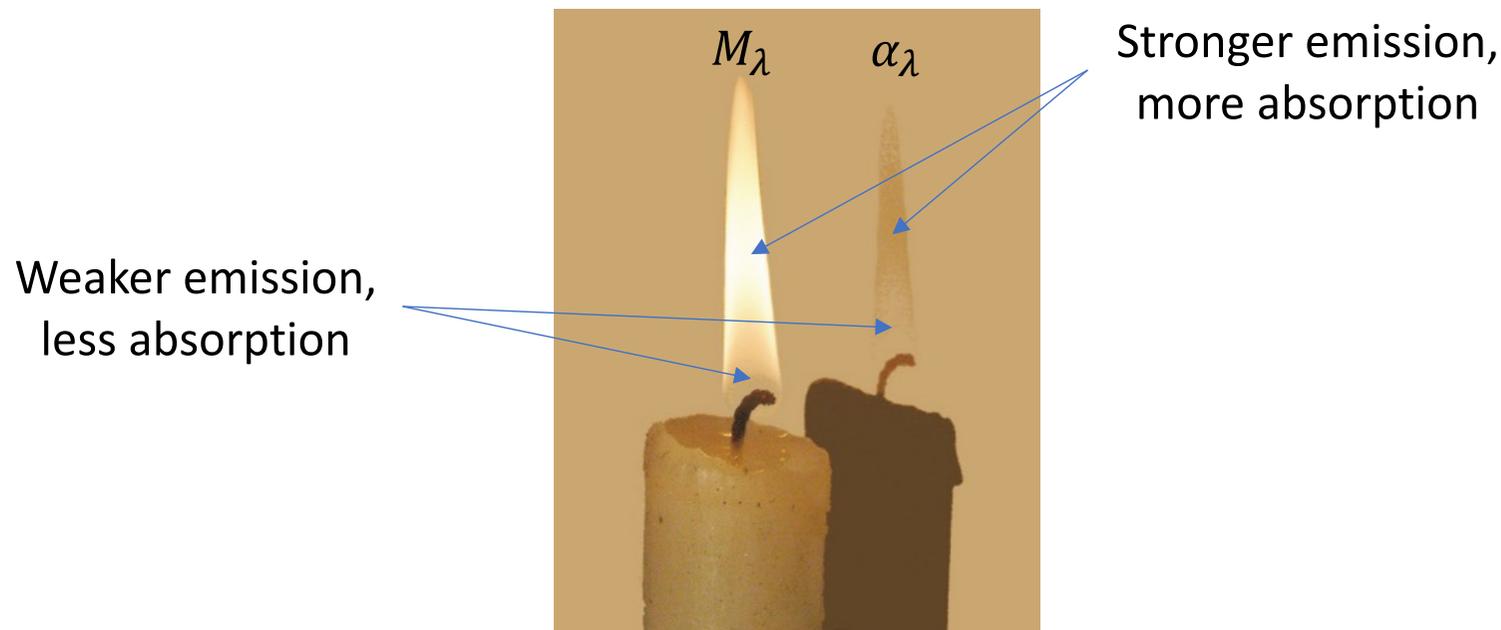
$M$  and  $\alpha$  strongly depend on the absolute temperature of the body!

# Kirchhoff's law

Bodies that emit more also absorb more. The ratio between radiant emittance and absorption coefficient is constant within a narrow range of wavelength ( $\lambda$ ):



$$\frac{M_{\lambda \text{ body1}}}{\alpha_{\lambda \text{ body1}}} = \frac{M_{\lambda \text{ body2}}}{\alpha_{\lambda \text{ body2}}} = \text{constant}$$



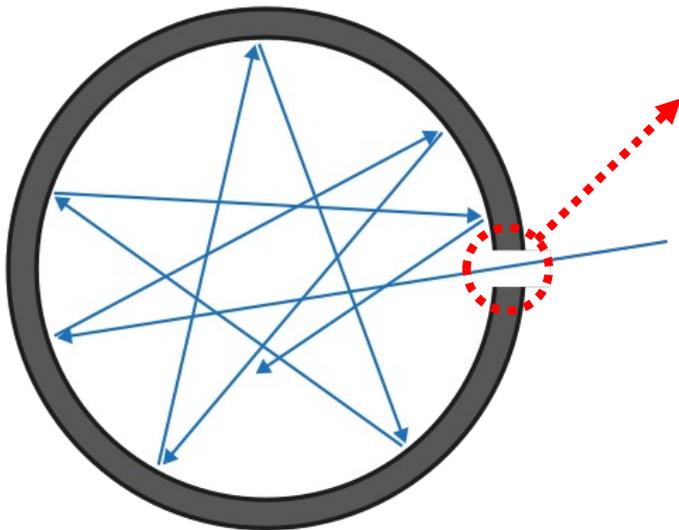
# The ideal black body

A theoretical body that is used as a model in the description of thermal radiation.  
It absorbs all radiation that falls on it:

$$\alpha_{black\ body} = 1$$

thus, we can calculate any real body's radiant emittance if we know its absorption coefficient ( $\alpha_{\lambda i}$ ):

$$M_{\lambda i} = \alpha_{\lambda i} M_{\lambda\ black\ body}$$



The hole on a dark cavity approximates an ideal black body

The radiant emittance of a black body depends strongly on the absolute temperature!

**Stefan's law:**

$$M = \sigma T^4$$

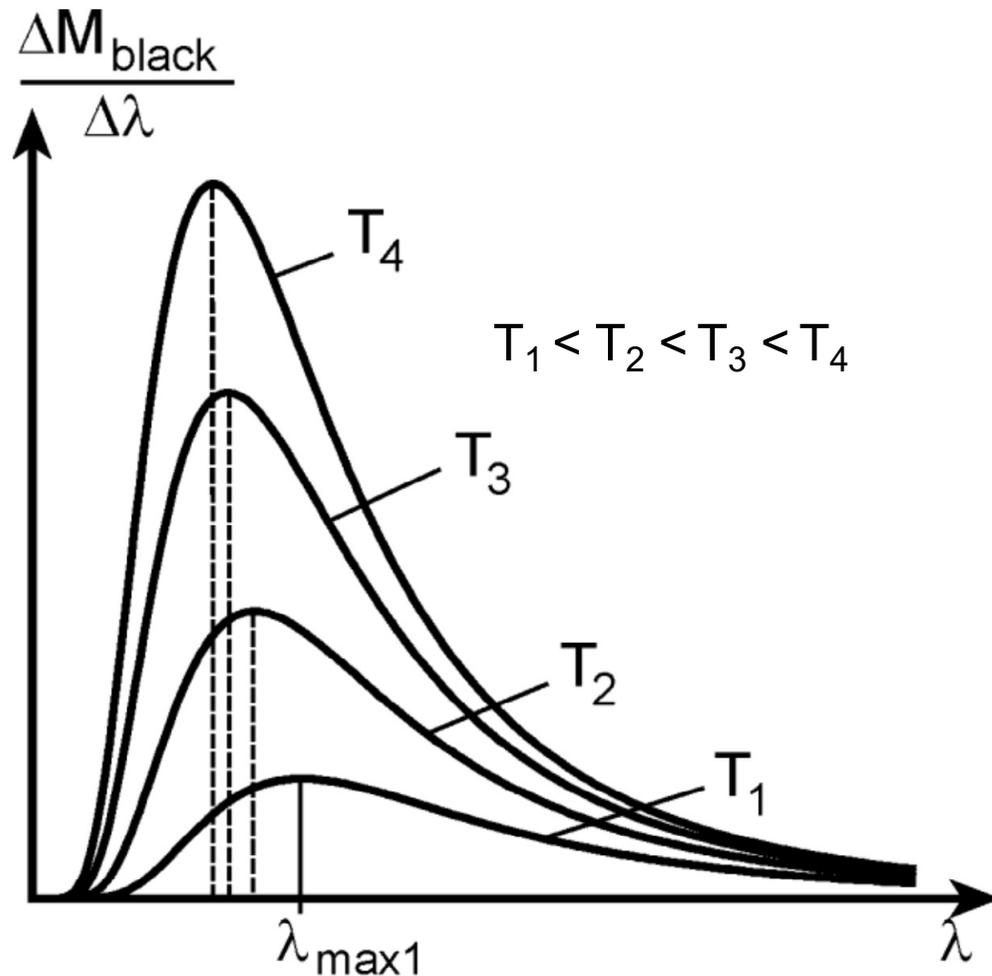
$$\sigma = 5.67 \cdot 10^{-8} \frac{W}{m^2 K^4}$$



Jozef Stefan

# The ideal black body

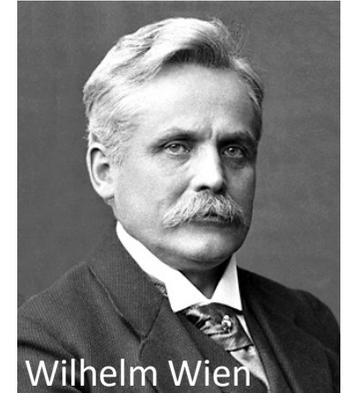
The emission spectrum of the black body is **continuous**.



**Wien's displacement law:**

$$\lambda_{\text{max}} T = \text{constant}$$

$$\text{constant} = 0.0029 \text{ m} \cdot \text{K}$$



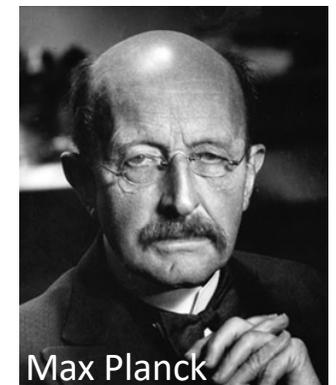
**Max Planck's radiation law:**

The radiation energy of a black body is emitted in discrete integer multiples of a fundamental "package", the quantum.

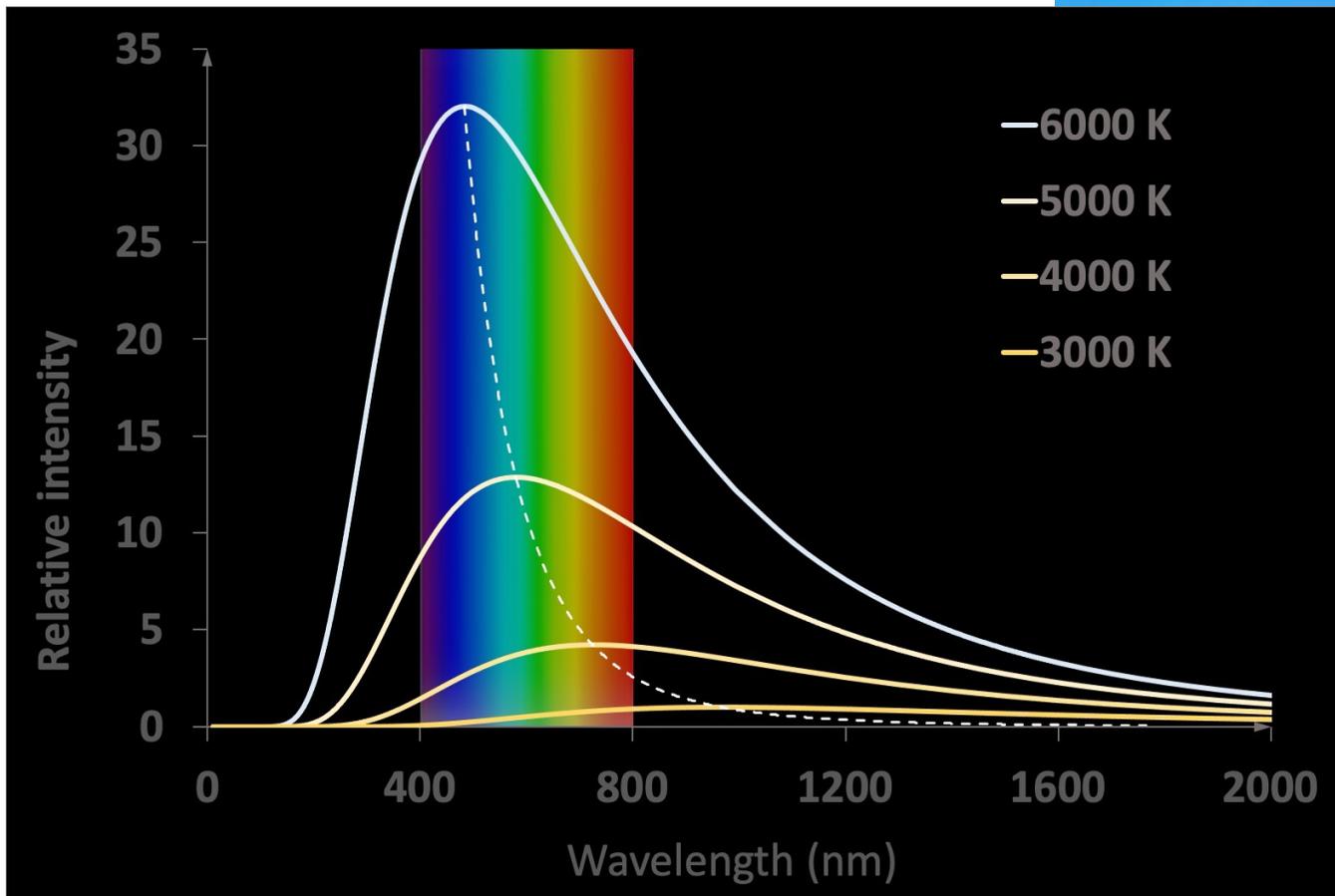
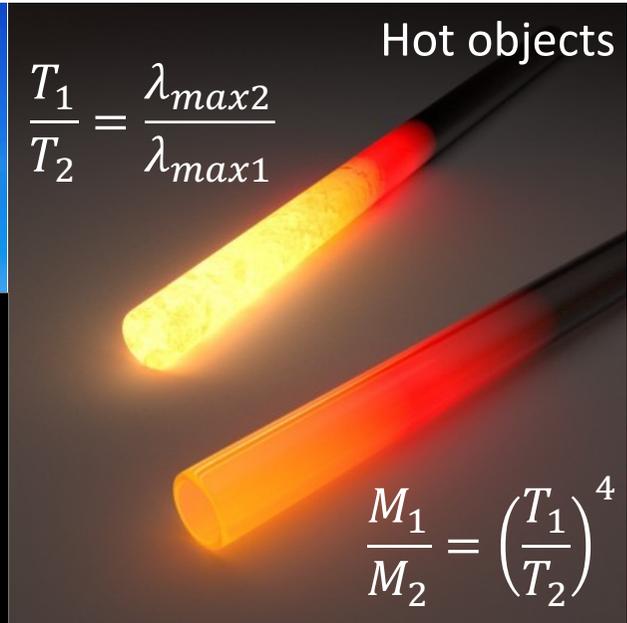
$$E = hf$$

$$h = 6.63 \cdot 10^{-34} \text{ J} \cdot \text{s}$$

$$M_{(f)} = \frac{2\pi h f^3}{c^2} \cdot \frac{1}{e^{\left(\frac{hf}{k_B T}\right)} - 1}$$



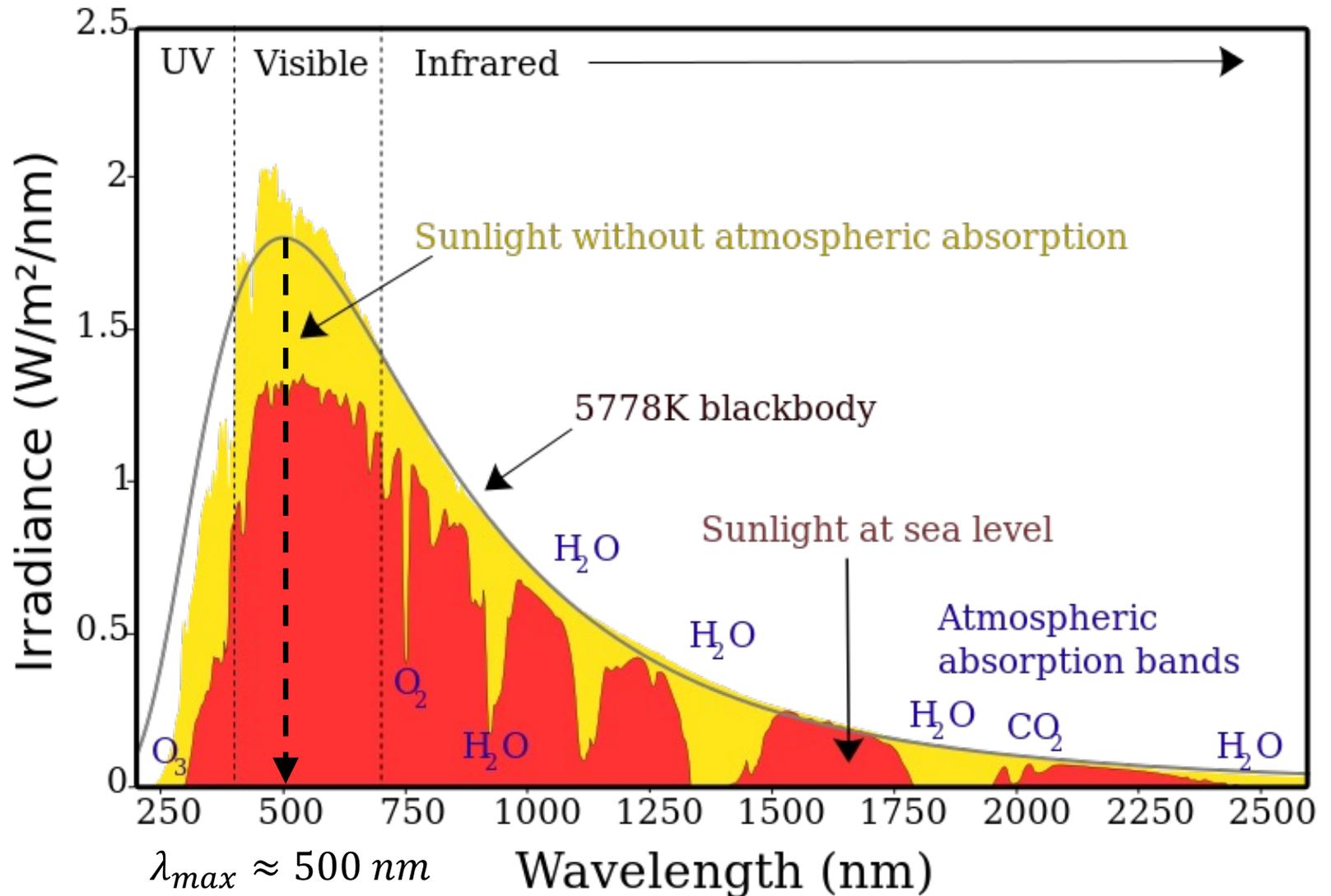
# Light sources based on thermal radiation



# Sun light



## Spectrum of Solar Radiation (Earth)



# Medical applications

## Telethermography

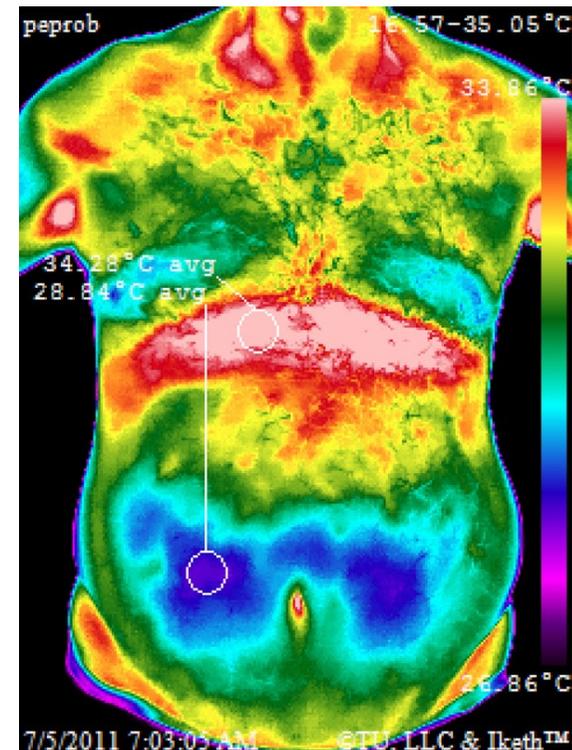
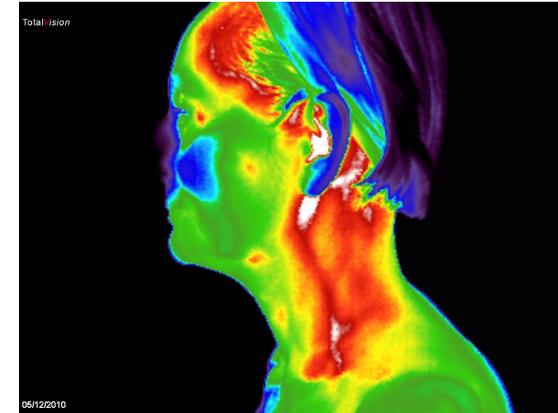
Measures the radiant emittance of thermal radiation of the human body.

Human body:  $\lambda_{max} \approx 10 \mu m$  (infrared)  
 $\alpha \approx 0.95$  (95% of ideal black body)

Radiant emittance strongly depends on temperature!



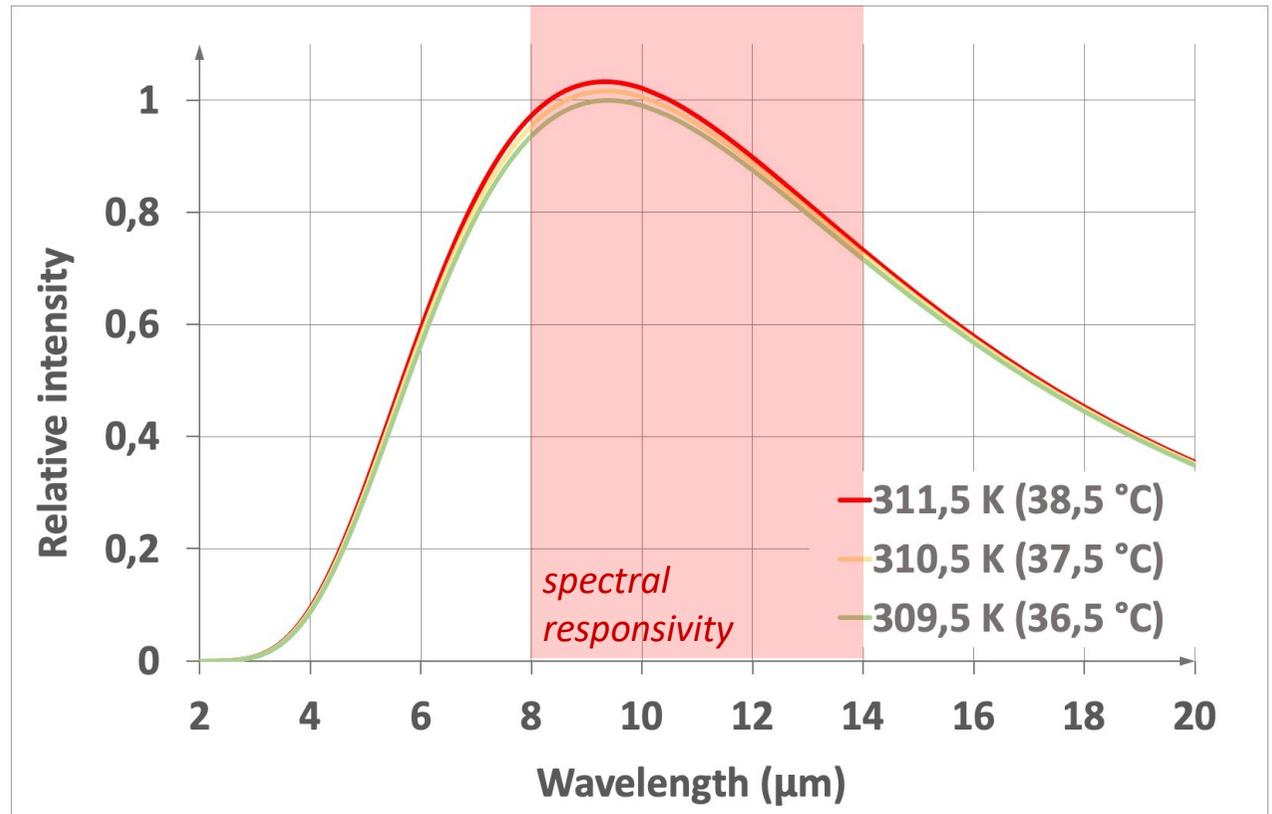
Diagnosis of inflammations, cancer, circulatory defects that cause the change in local temperature.



# Medical applications

## Non-contact thermometry

Typically measures the total radiant intensity between 8-14  $\mu\text{m}$  (*spectral responsivity*).  
1  $^{\circ}\text{C}$  of temperature difference will increase the radiante emittance by  $\sim 1.5\%$  !



# Luminescence

Spontaneous emission of a photon due to the relaxation of an excited electron.

Phases of luminescence:

- Absorption of external energy
- Excitation
- Emission of energy in the form of electromagnetic radiation

## Types of luminescence

Type of excitation	Name	Example
Light	Photoluminescence	Fluorescent lamp
Electric	Electroluminescence	Na-lamp
Radioactive	Radioluminescence	NaI (Tl) (scintillator)
Mechanical	Triboluminescence	(Percussion, friction)
Biochemical	Bioluminescence	firefly
Thermal	Thermoluminescence	CaSO <sub>4</sub> (Dy) (dosimeter)

Type of relaxation	Name	Example
$S_1 - S_0$ (fast)	Fluorescence	Fluorescein
$T_1 - S_0$ (slow)	Phosphorescence	Phosphorous

Fluorescent dyes



Bioluminescence



LED



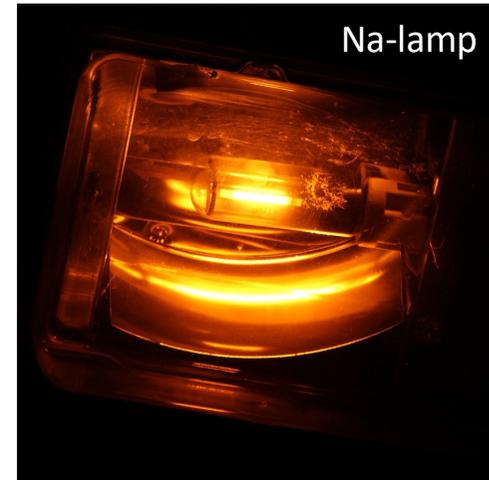
Phosphorescent dial



Radioluminescence - scintillation

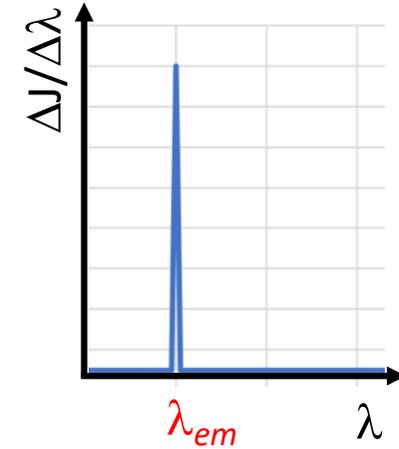
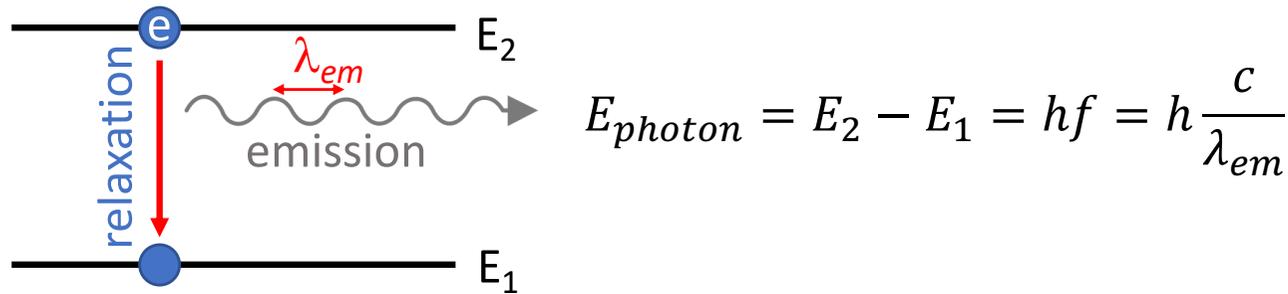


Na-lamp

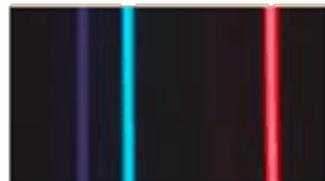
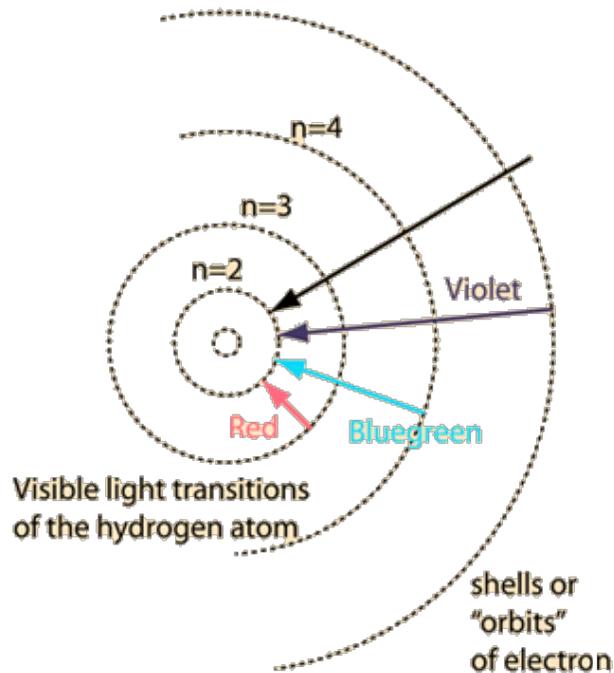
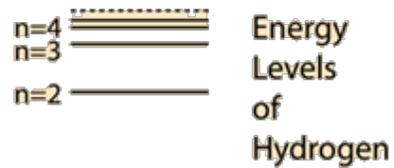


# Luminescence of atoms

Luminescence emission of atoms has a **line spectrum**.

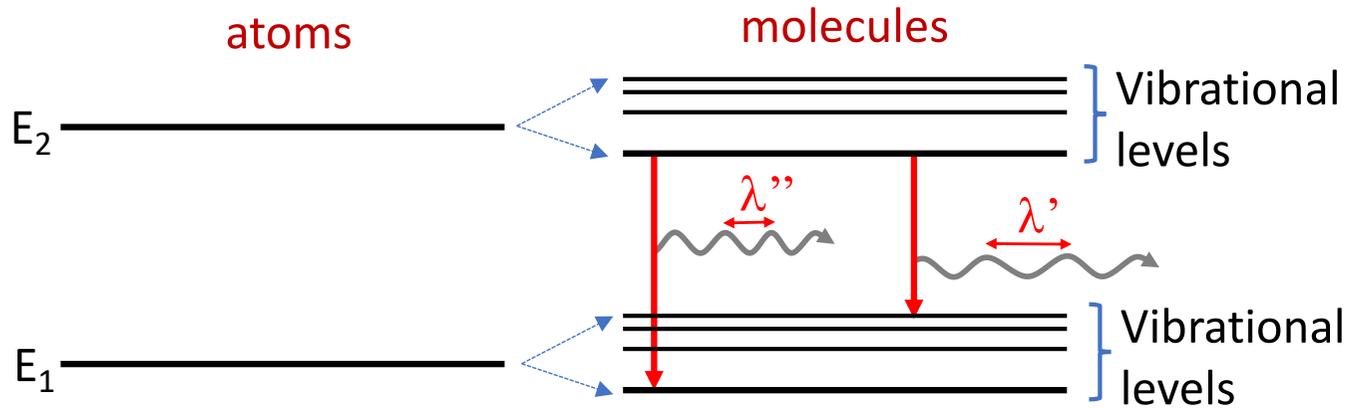


Flame test



# Luminescence of molecules

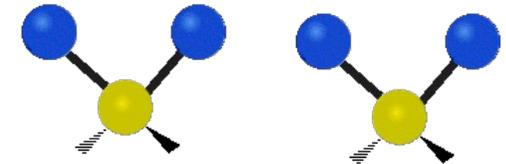
Discrete energy levels split into **vibrational levels**.



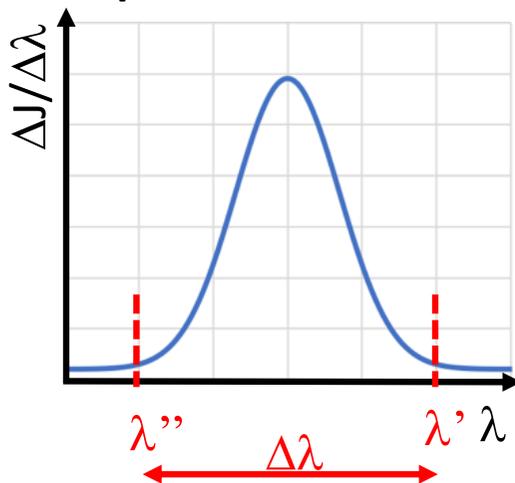
The energy of a molecule is the sum of its electronic-, vibrational, and rotational transition energies. :

$$E_{total} = E_e + E_v + E_r$$

Molecular vibrations:  
(2 examples)



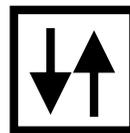
Luminescence of molecules has **band spectrum**



## Spin states of excited electrons

### Singlet state (S)

Sum of spin quantum numbers is  $S = 0$  (+1/2, -1/2)



### Triplet state (T)

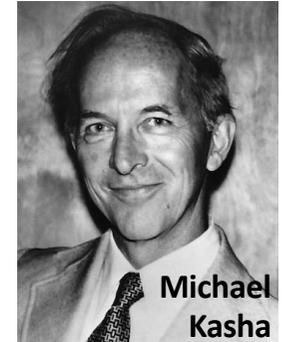
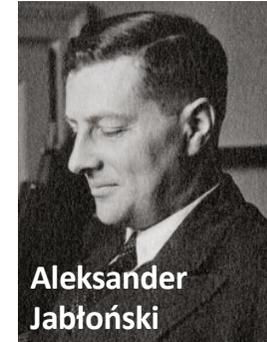
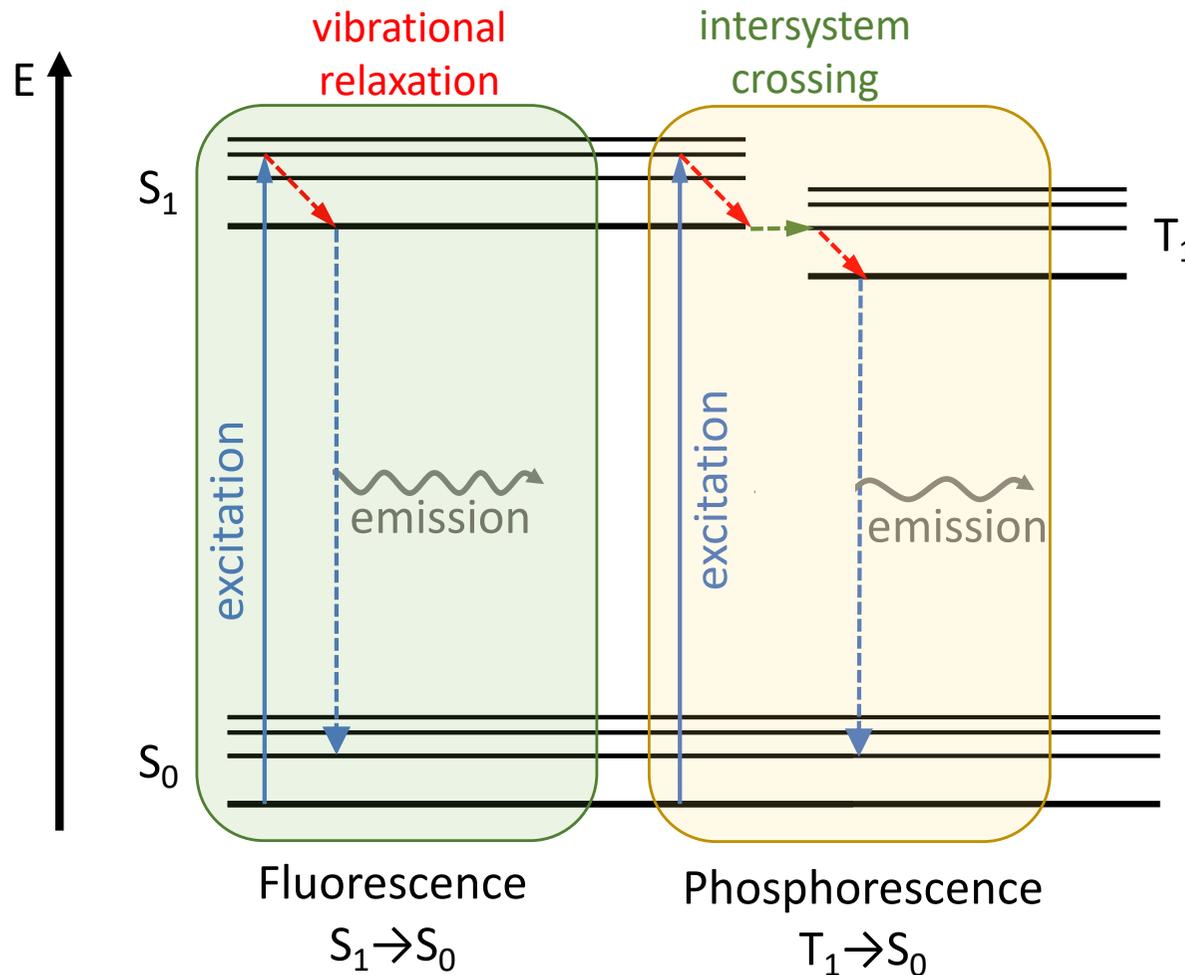
Sum of spin quantum numbers is  $S = 1$  (+1/2, +1/2)



Magnetic moment of spin state:  $2S + 1$

# Luminescence emission of molecules

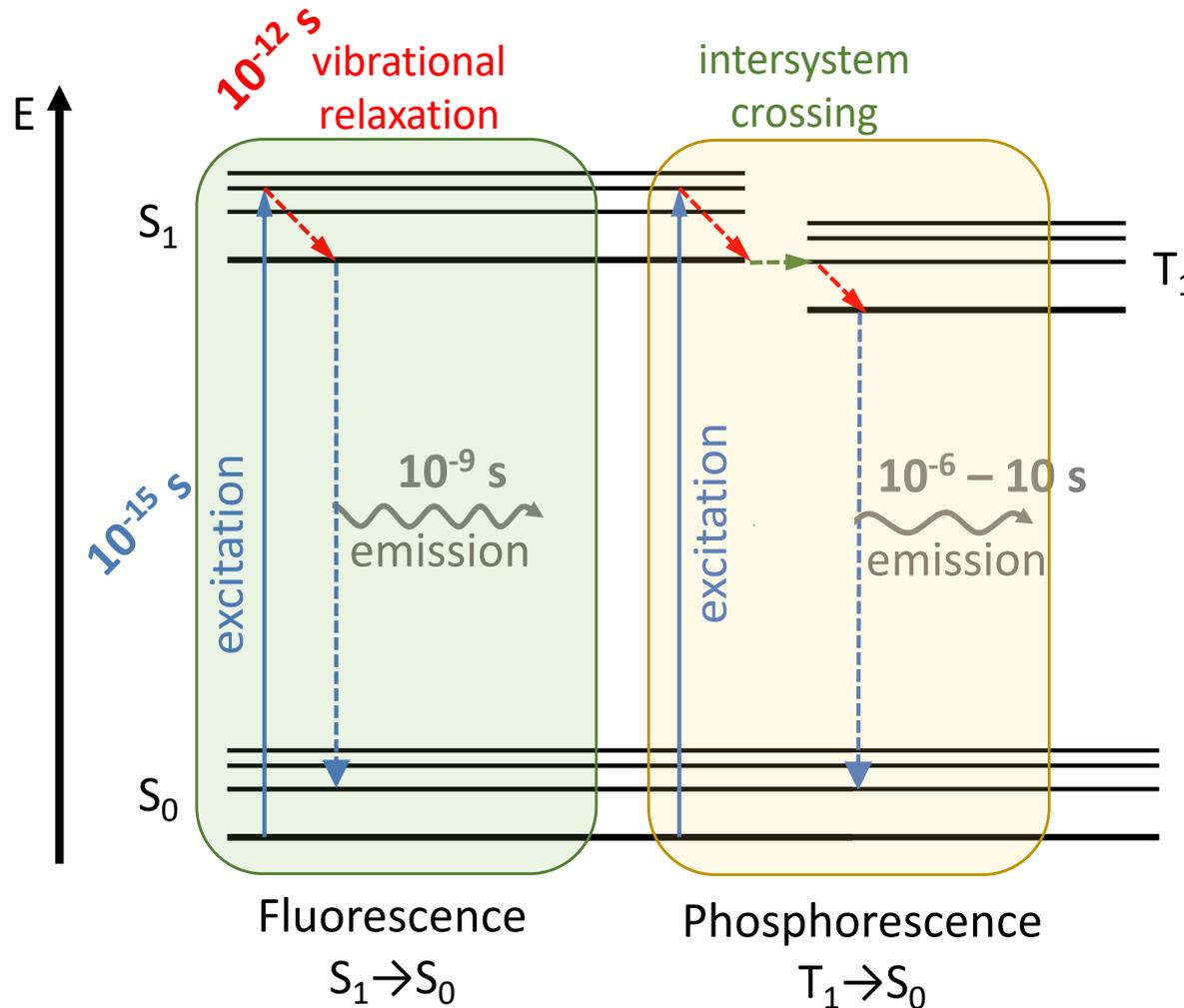
## Jabłoński diagram



**Kasha's rule:** The excited molecule first reaches the lowest vibrational level of  $S_1$  and photon emission occurs always from this state to any vibrational level of the ground ( $S_0$ ) state.

# Luminescence emission of molecules

## Jabłoński diagram



### Lifetimes ( $\tau$ )

Excitation:  
femtosecond ( $10^{-15}$  s)

Vibrational relaxation:  
picosecond ( $10^{-12}$  s)

Fluorescence emission:  
nanosecond ( $10^{-9}$  s)

Phosphorescence emission:  
microsecond to seconds  
( $10^{-6} - 10$  s)

# Luminescence emission of molecules

Quantum yield ( $Q_F$ ):

$$Q_F = \frac{k_f}{k_f + k_{nr}} = \frac{\text{number of photons emitted}}{\text{number of photons absorbed}}$$

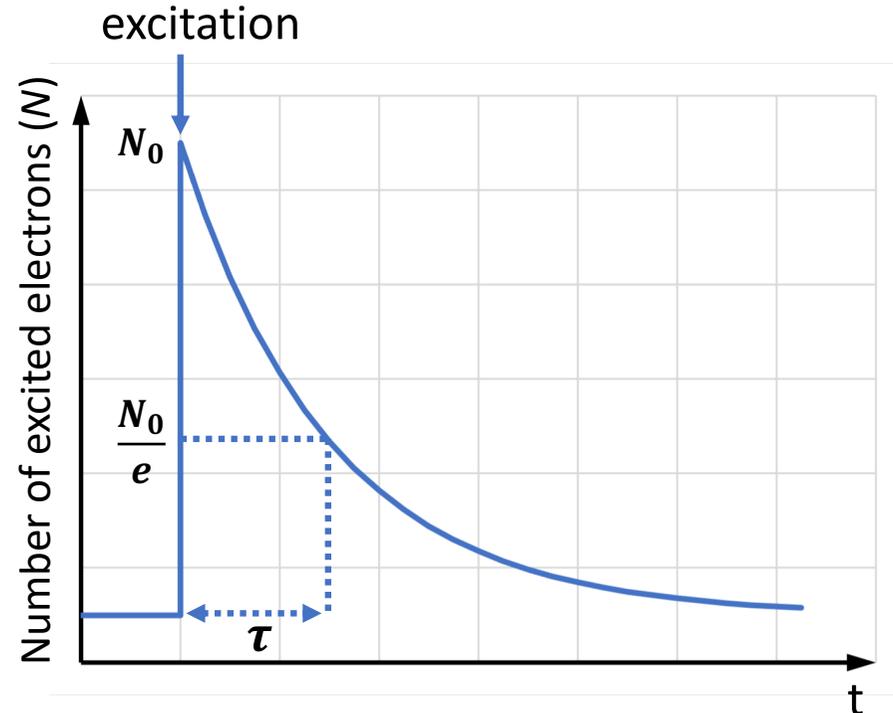
Lifetime ( $\tau$ ):

$$N = N_0 \cdot e^{-\frac{t}{\tau}}$$

$$\tau = \frac{1}{k_f + k_{nr}}$$

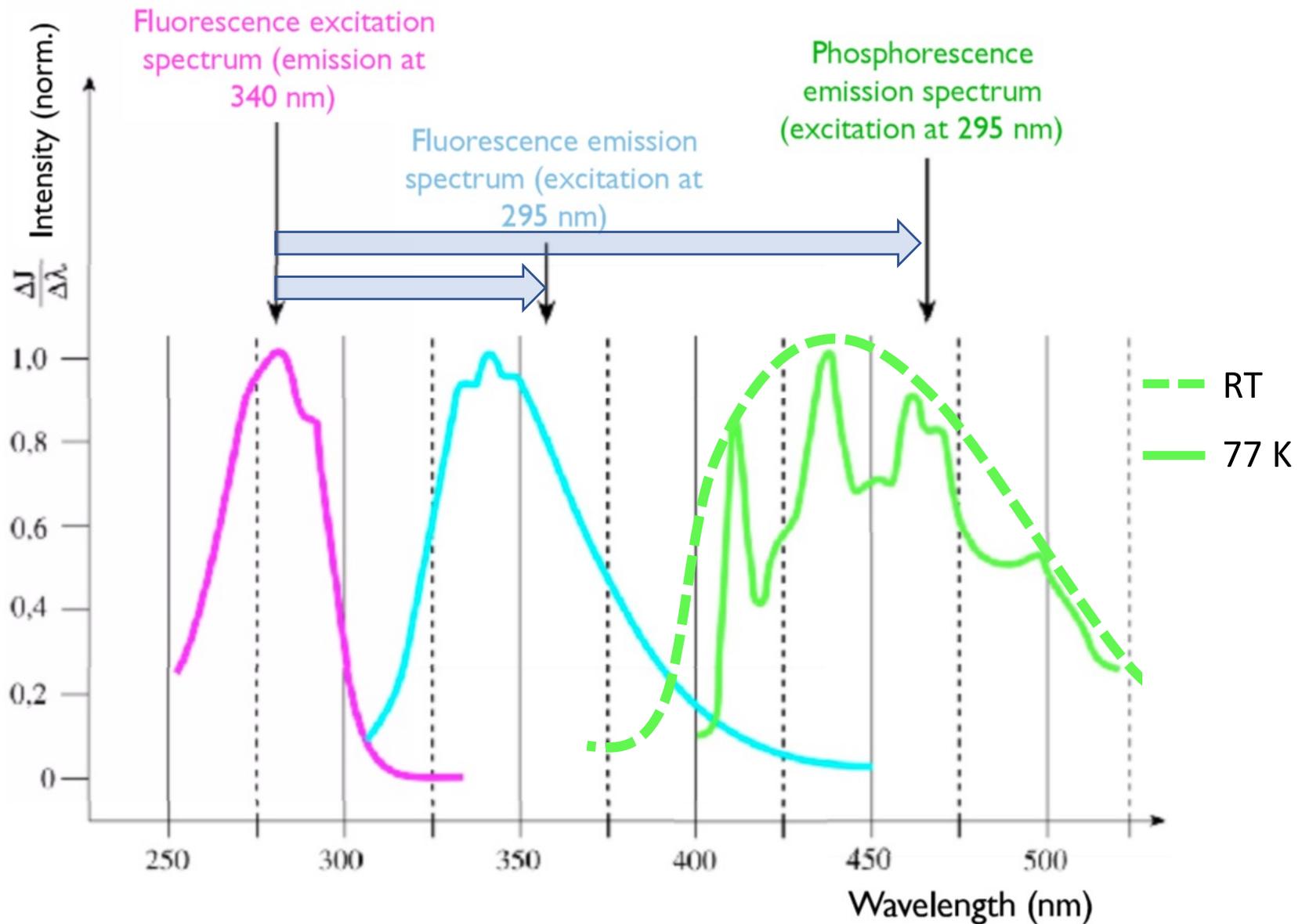
$k_f$ : Rate of photon producing transitions

$k_{nr}$ : Rate of non-radiative transitions



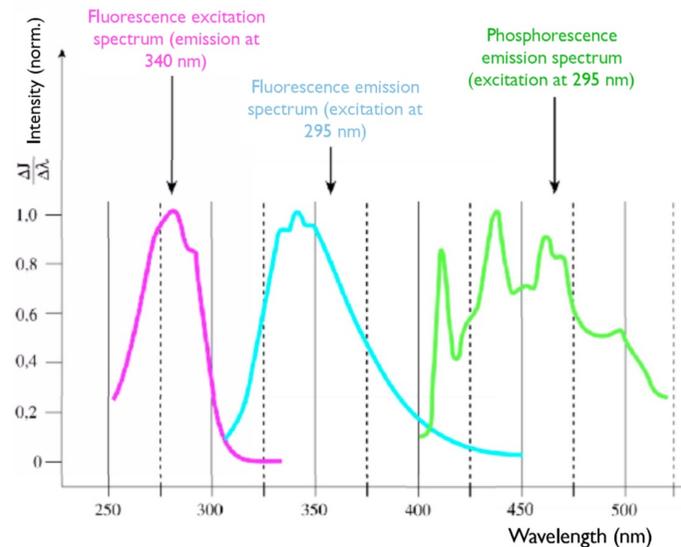
# Stokes shift

The shift between the emission and excitation spectra



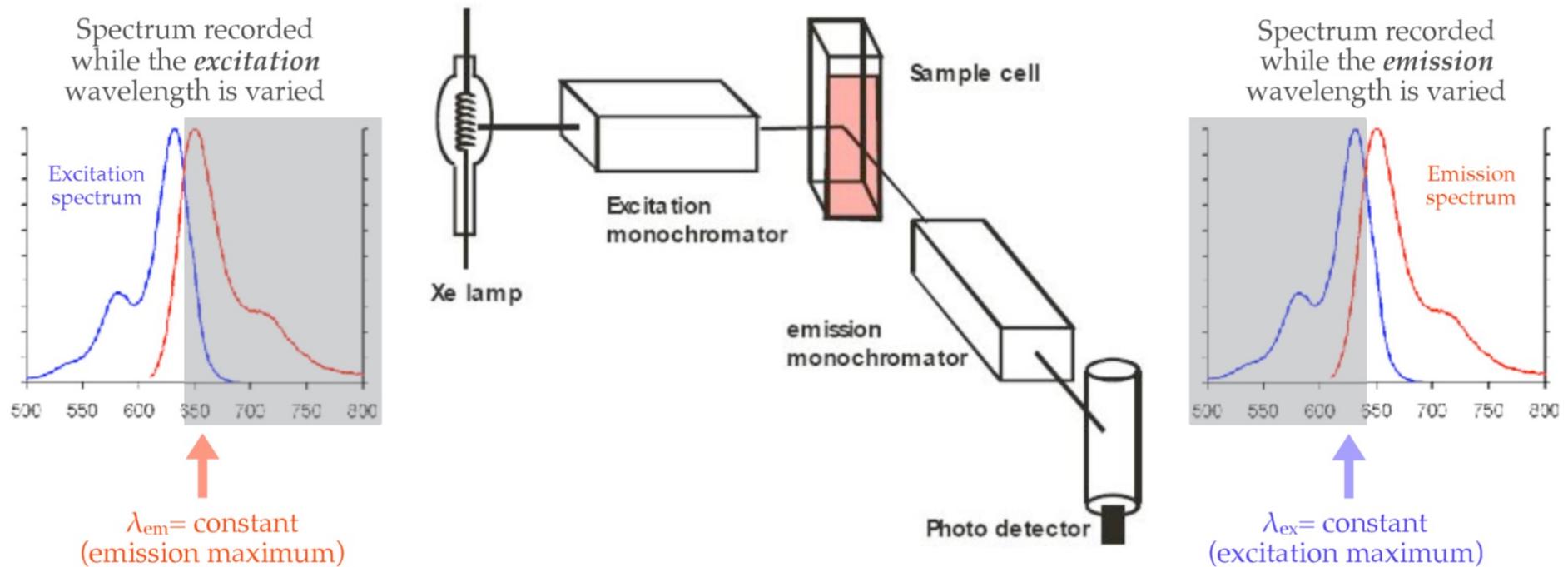
# Comparing emission processes

	Fluorescence	Phosphorescence
Relaxation	From singlet state $S_1 \rightarrow S_0$	From triplet state $T_1 \rightarrow S_0$
Lifetime	Nanoseconds	From microseconds up to seconds
Stokes shift	Smaller	Larger (because $T_1$ is at lower energy than $S_1$ )



# Measurement of luminescence

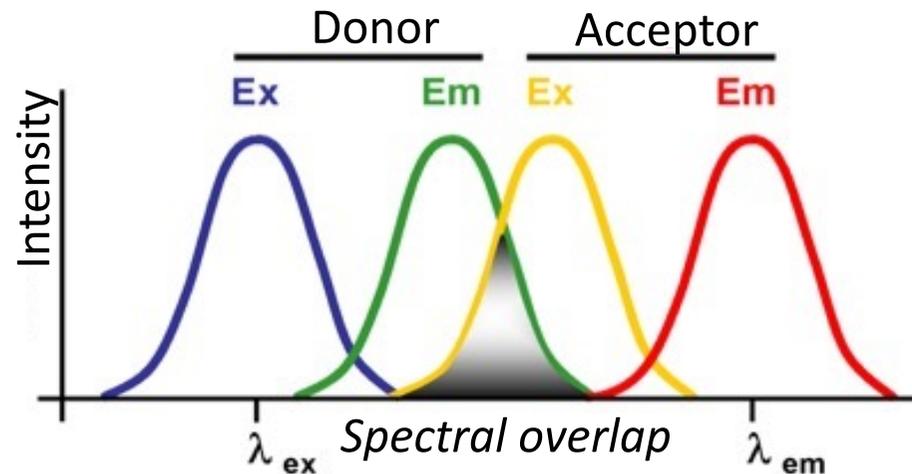
## Luminescence spectrofluorimeter



# Applications - FRET

## Förster Resonance Energy Transfer

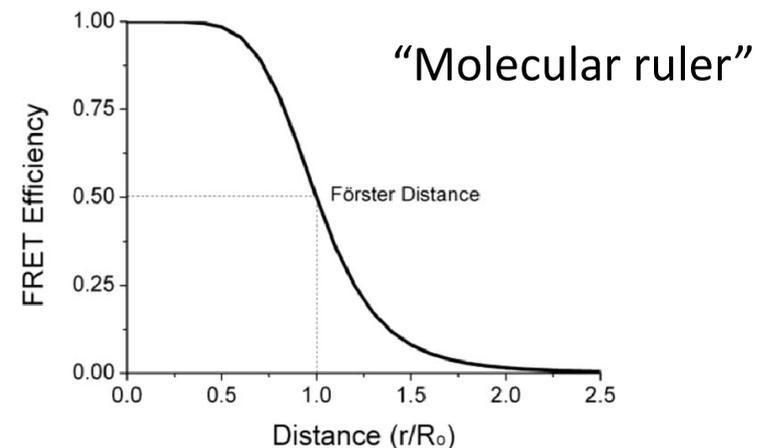
Energy transfers from donor without emission to acceptor in dipole-dipole interactions. Requires spectral overlap between donor emission and acceptor absorption.



FRET efficiency or quantum yield ( $E$ ):

$$E = \frac{1}{1 + (r/r_0)^6}$$

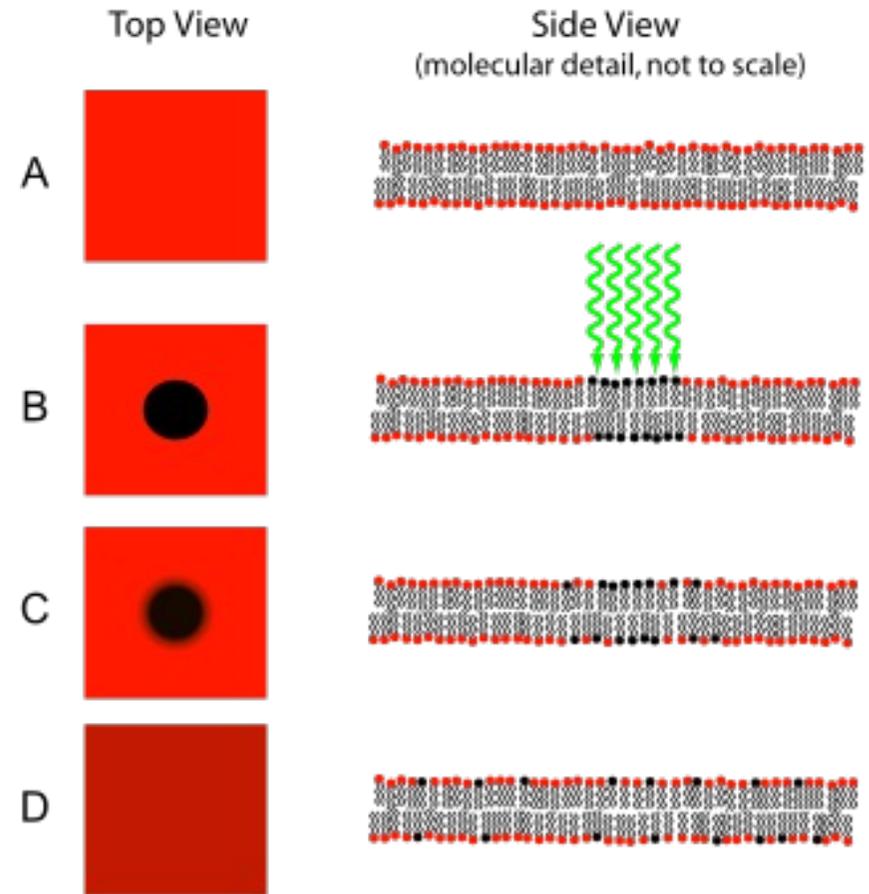
$r_0$ : Förster distance



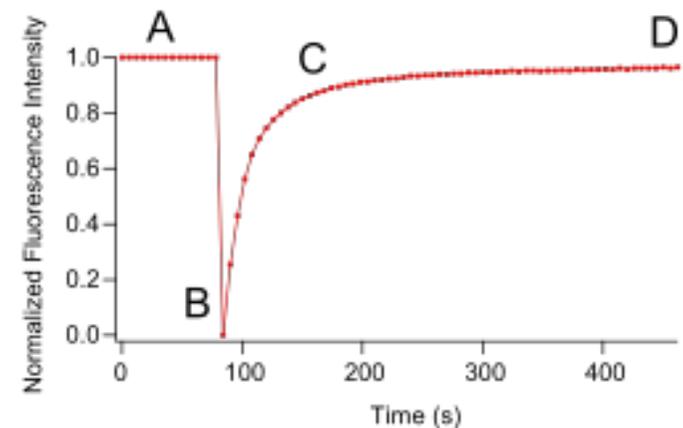
# Applications - FRAP

## Fluorescence Recovery After Photobleaching

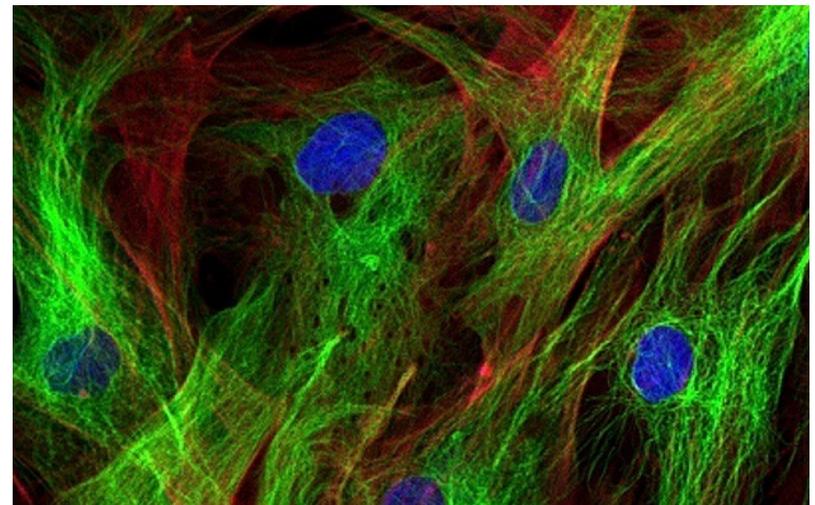
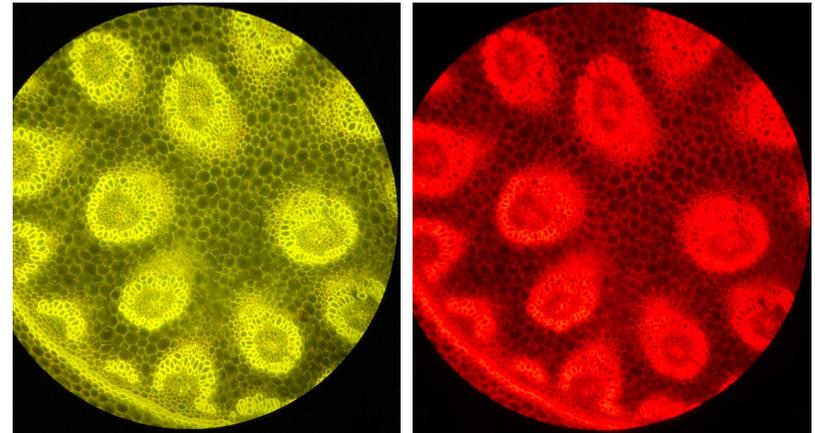
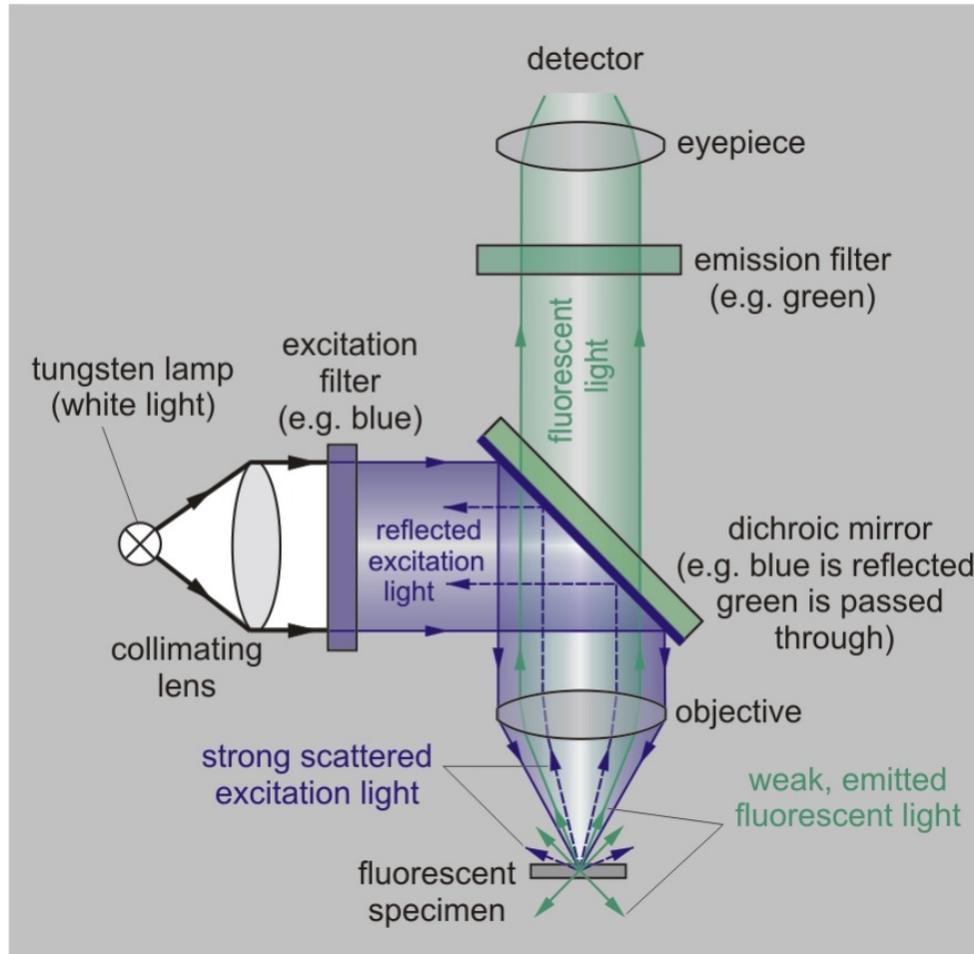
Photobleaching: The permanent loss of fluorescence due to photochemical reactions.



Lateral diffusion  
of lipid molecules

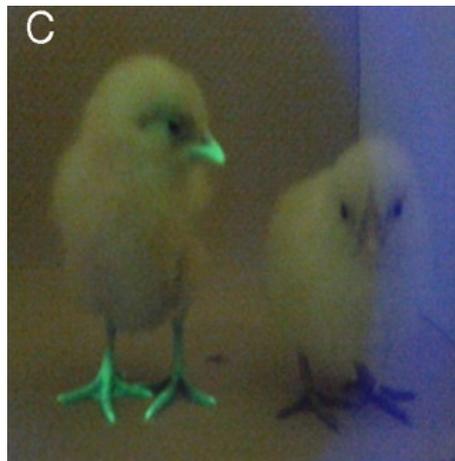


# Fluorescence microscope



# Biomedical applications

Fluorescent transgenic animals



Fluorescence guided surgery

